

Data inversion techniques

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ISMAR-CNR

IR0000032 – ITINERIS, Italian Integrated Environmental Research Infrastructures System
(D.D. n. 130/2022 - CUP B53C22002150006) Funded by EU - Next Generation EU PNRR-
Mission 4 “Education and Research” - Component 2: “From research to business” - Investment
3.1: “Fund for the realisation of an integrated system of research and innovation infrastructures”



Data Inversion Techniques: Outline

🌐 Information 'Space' dimensions

🌐 Inversion Approaches

Data Inversion Techniques: Outline

 **Information 'Space' dimensions**

 Inversion Approaches

INVERSION OF THE REMOTELY SENSED SIGNAL

The measured radiation L can be considered as a vector in a multidimensional space:

$$L = f \left[\lambda, s_{x,y,z}, t, (\theta_s, \theta_v, \varphi), (I, Q, U, V) \right]$$

where

λ = wavelength

$s_{x,y,z}$ = x, y, z position in the space (e.g. in the image)

t = time

$\theta_s, \theta_v, \varphi$ = set of angles that define the observation-illumination geometry (i.e the path: radiation source-observed object-instrument)

I, Q, U, V = Polarization of the measured radiation

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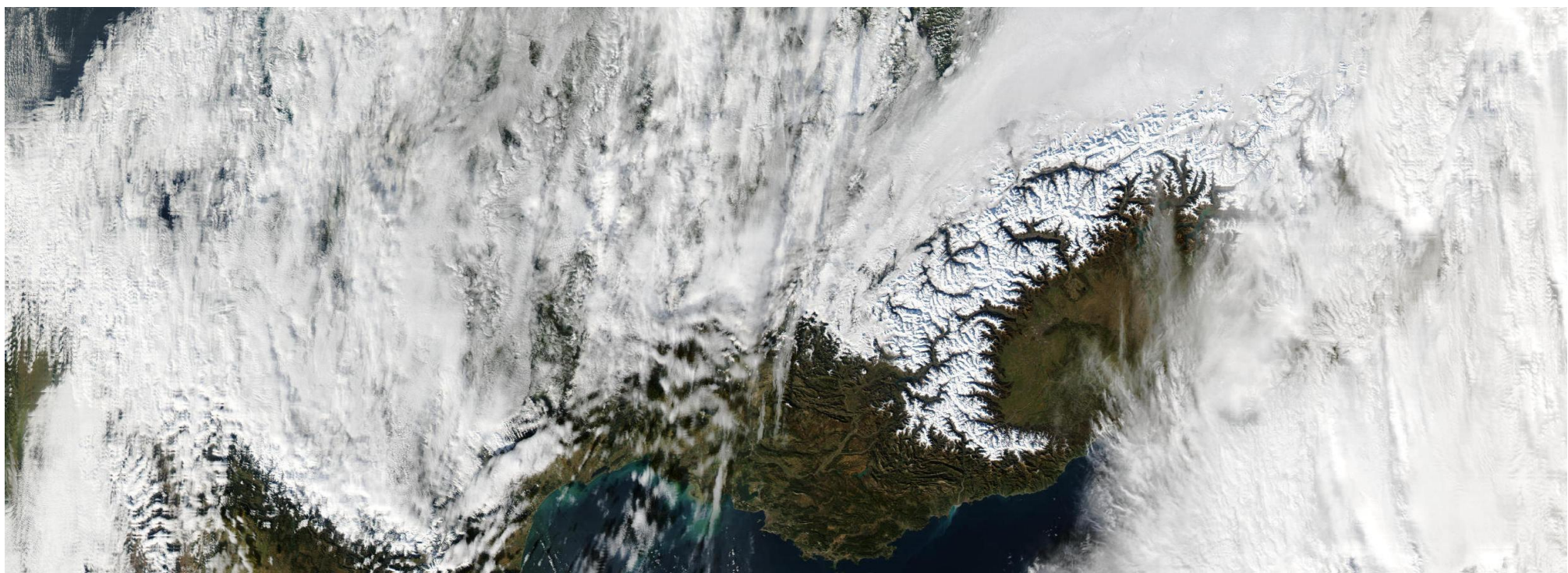
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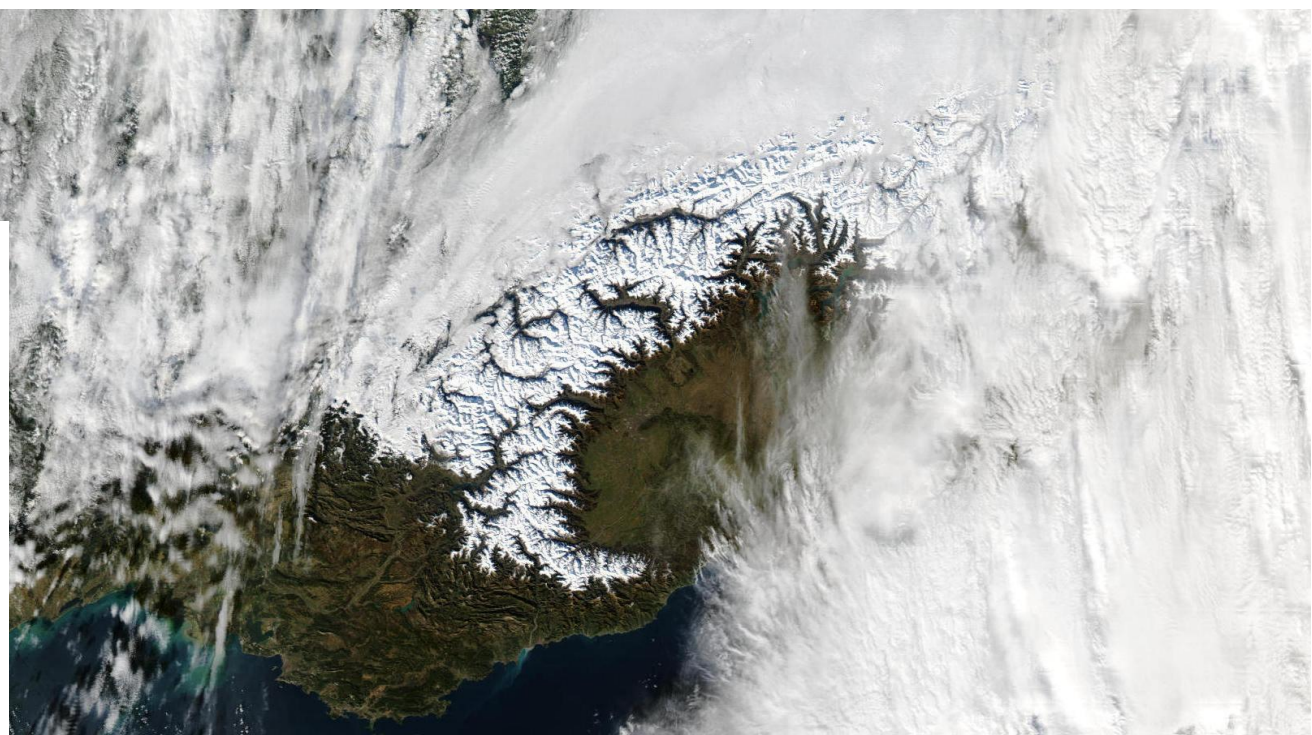
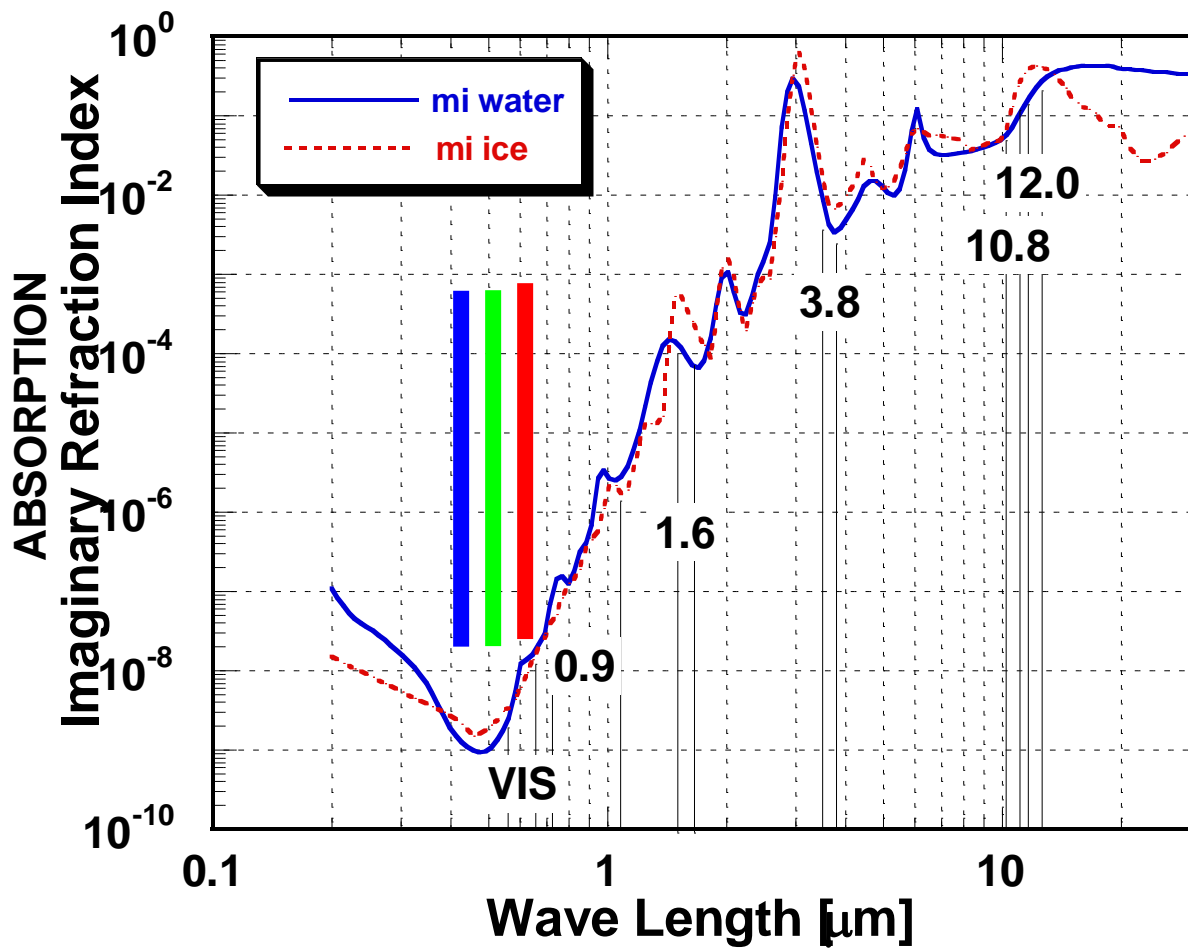
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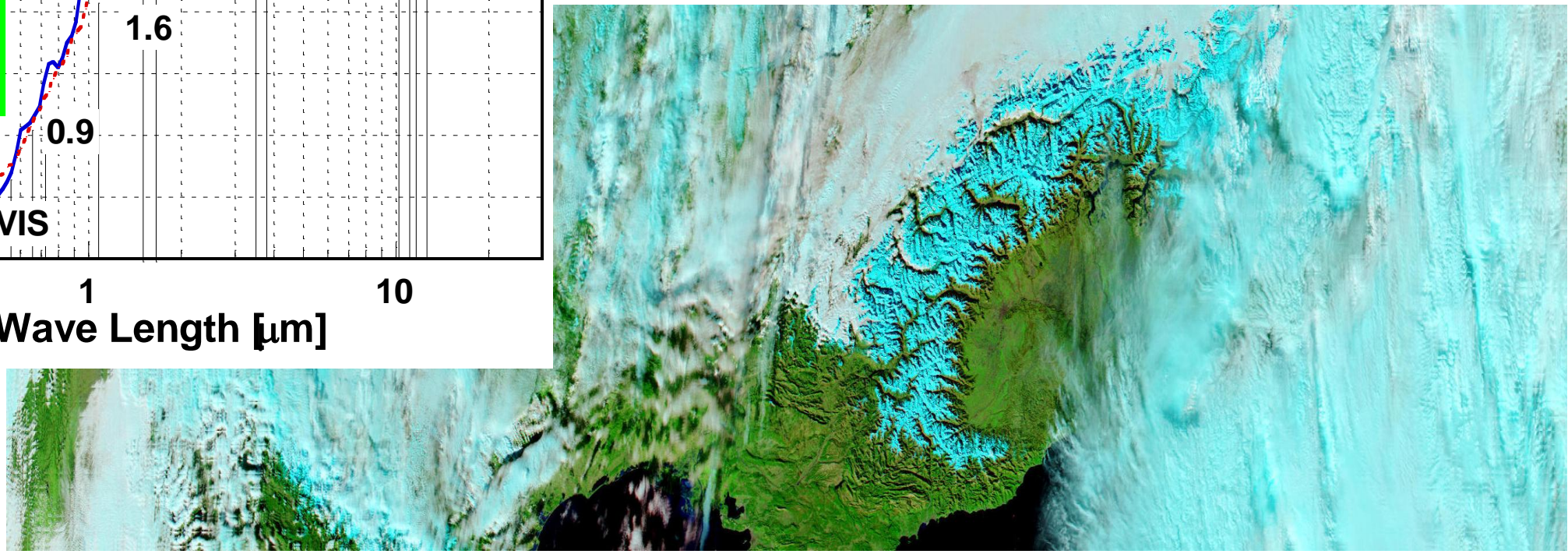
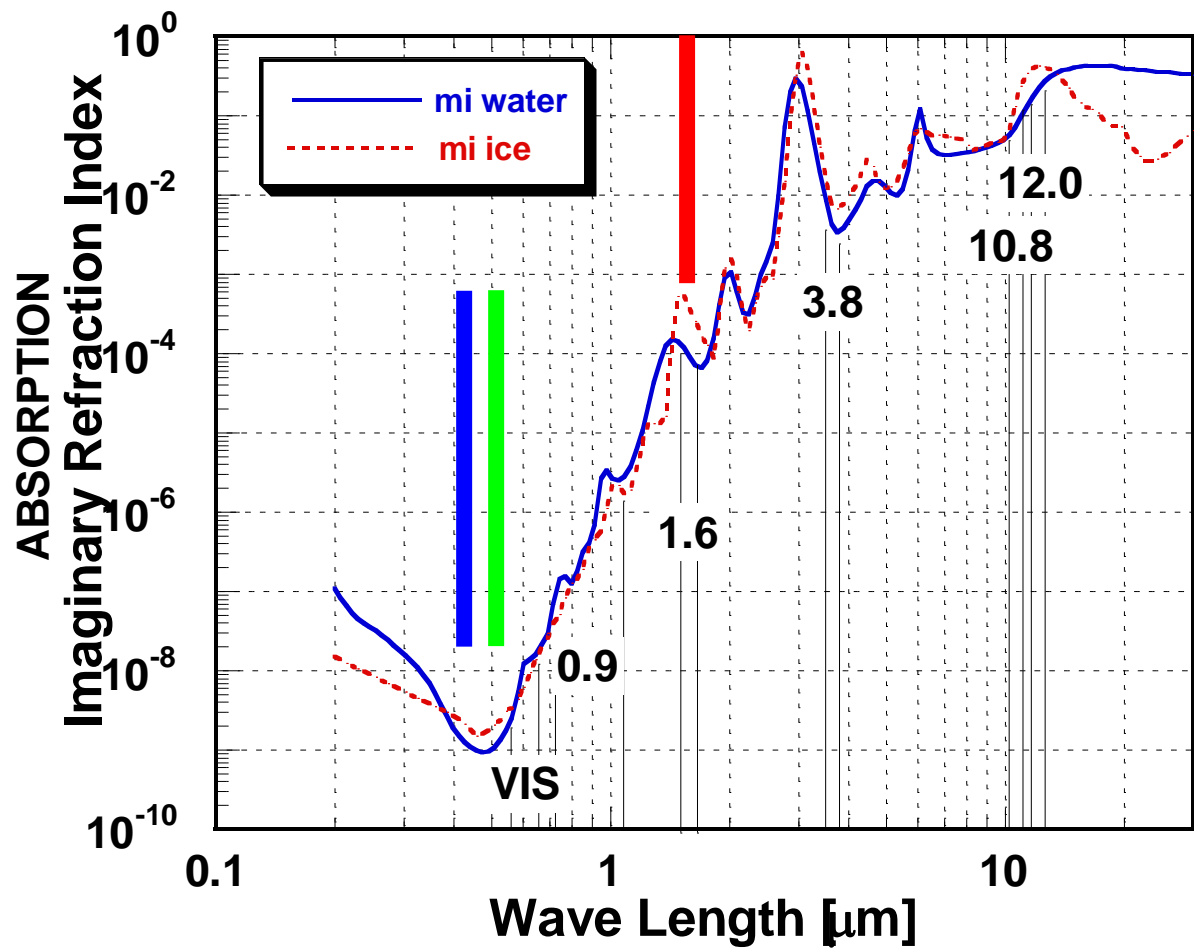
t = time

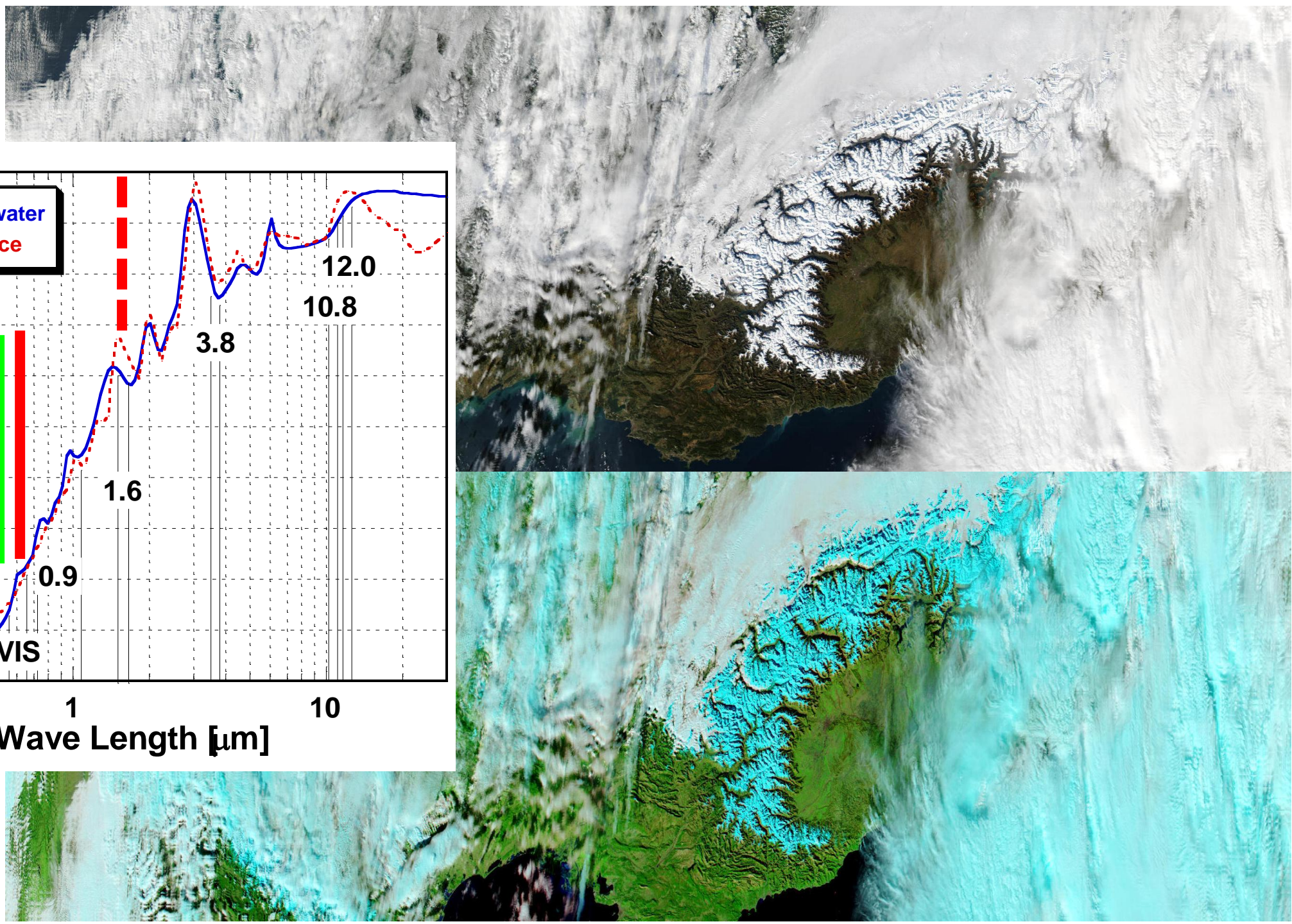
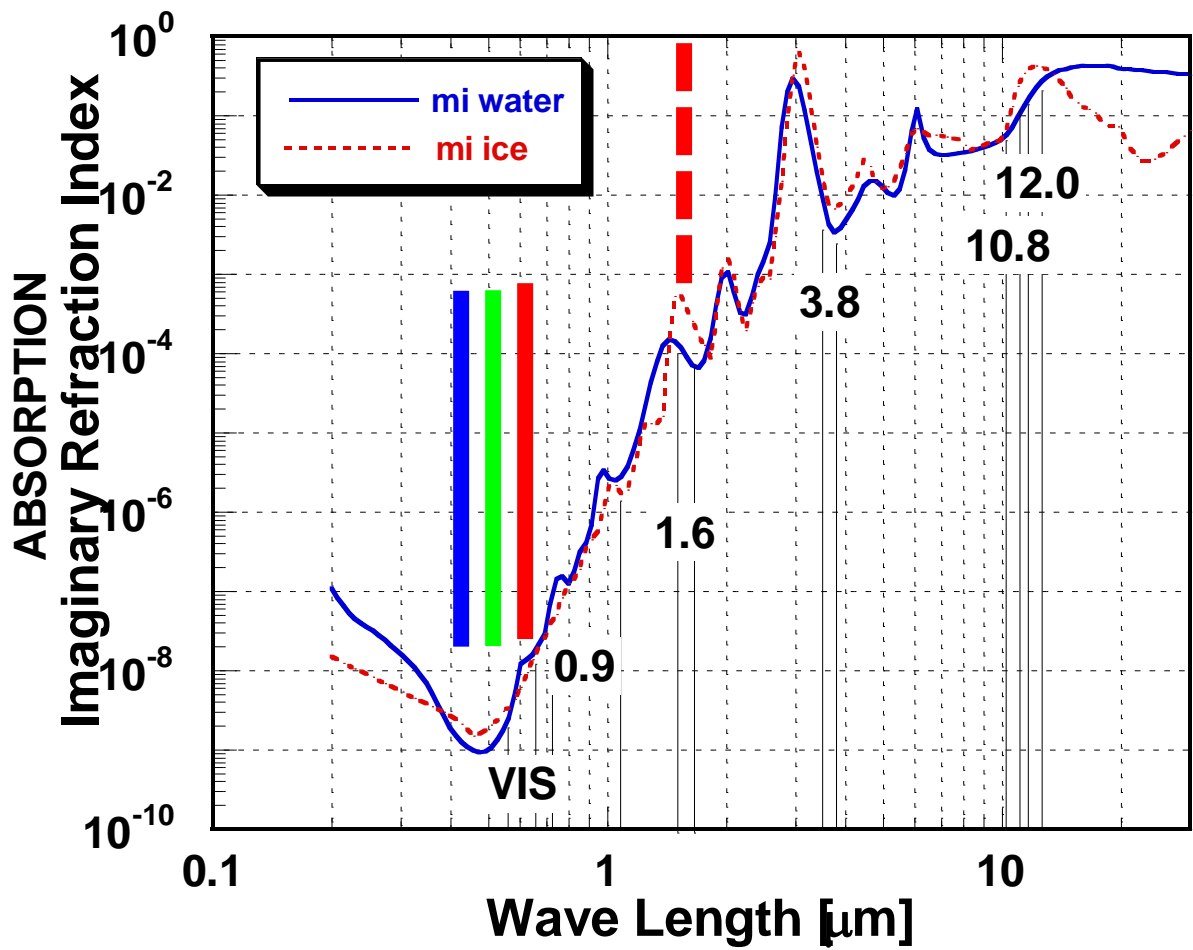
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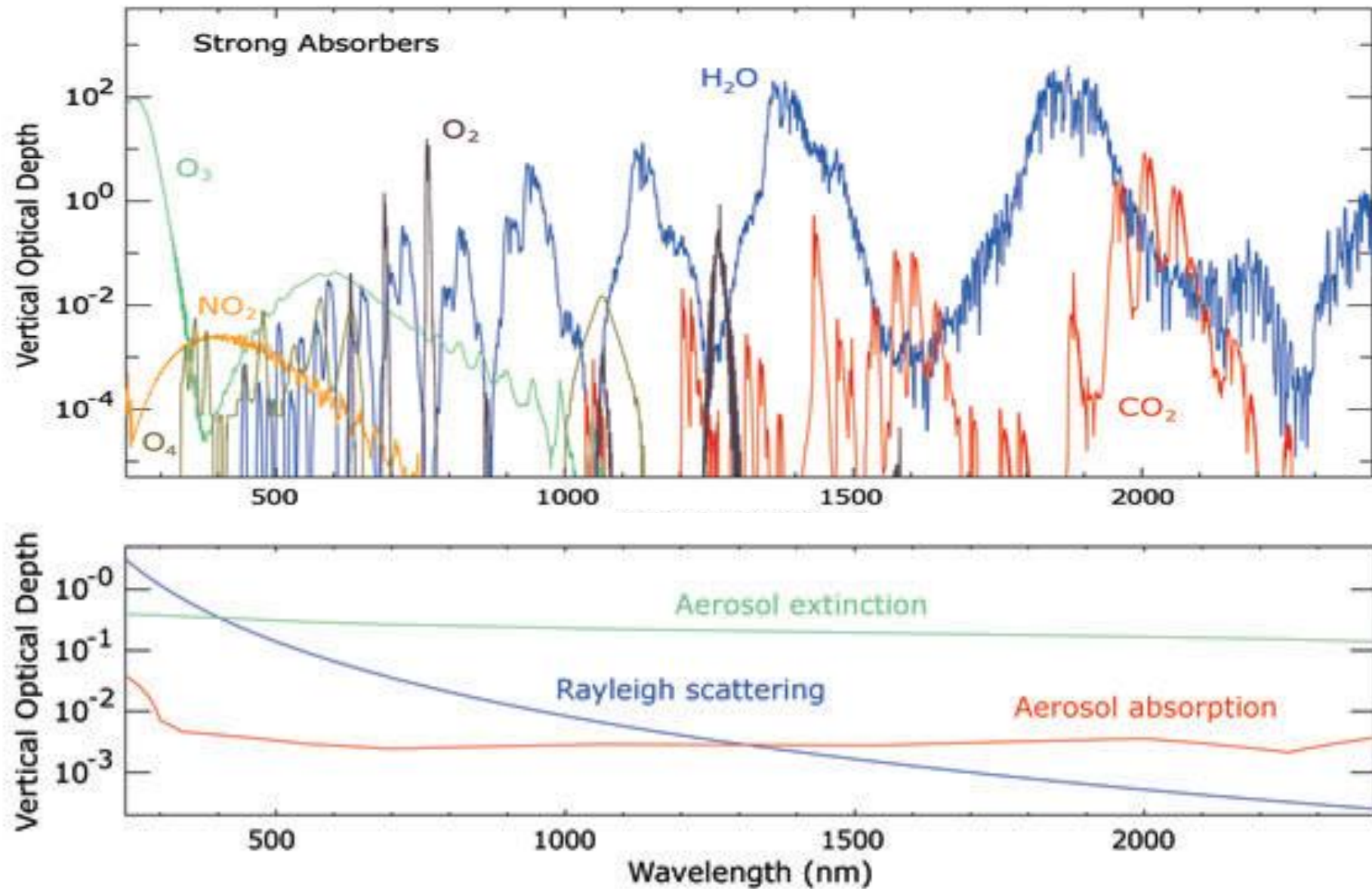
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Examples of gas absorption spectra (upper panel) and scattering extinction (lower panel) in the solar (UV/VIS/SWIR) range: **scattering spectral dependence smoother than gas absorption one**

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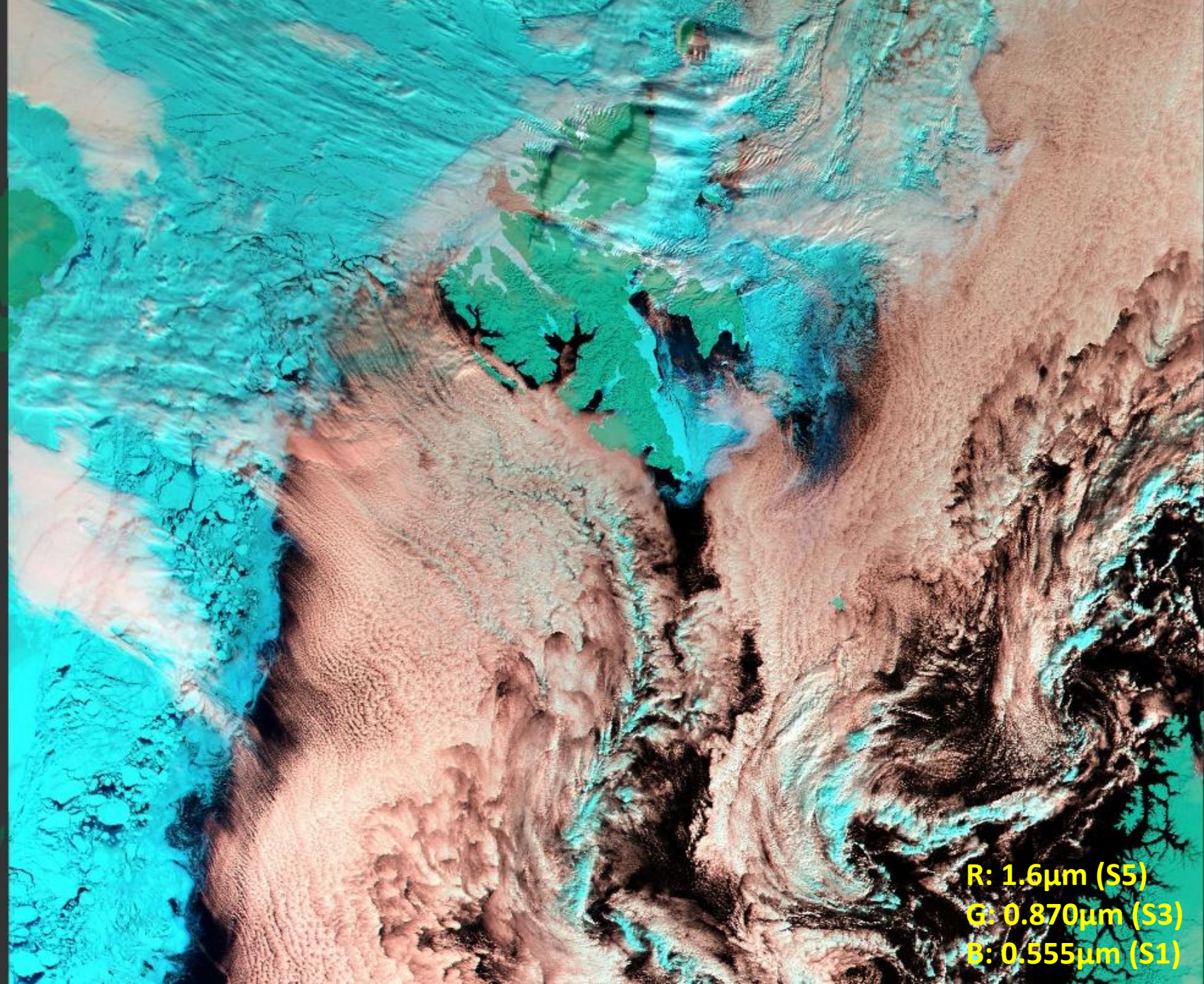
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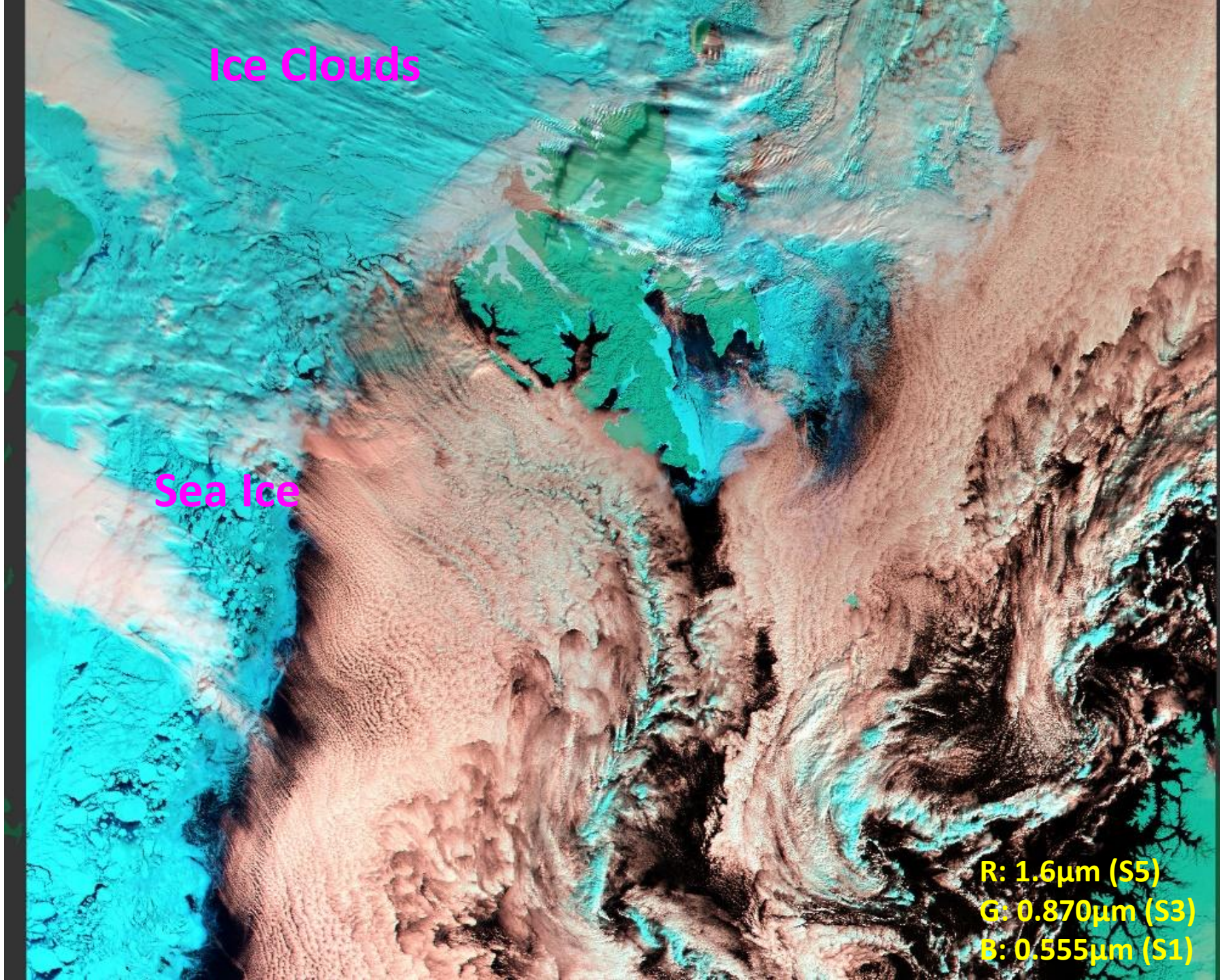
Can you
distinguish **Sea Ice**
from **Ice Clouds**?



Ice Clouds

Sea Ice

R: 1.6 μ m (S5)
G: 0.870 μ m (S3)
B: 0.555 μ m (S1)



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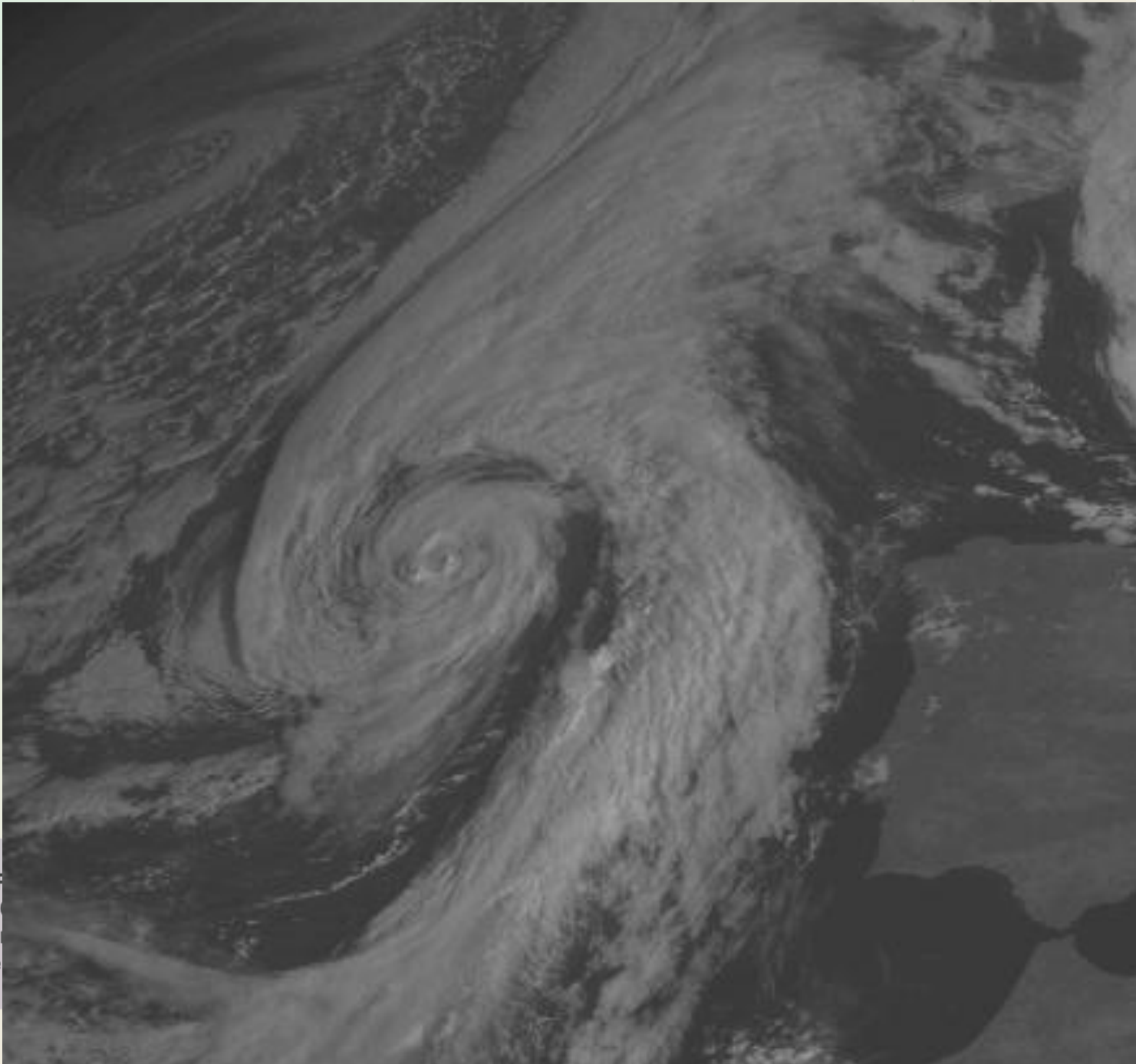
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Mission 4 “Education
3.1: “Fund for the re



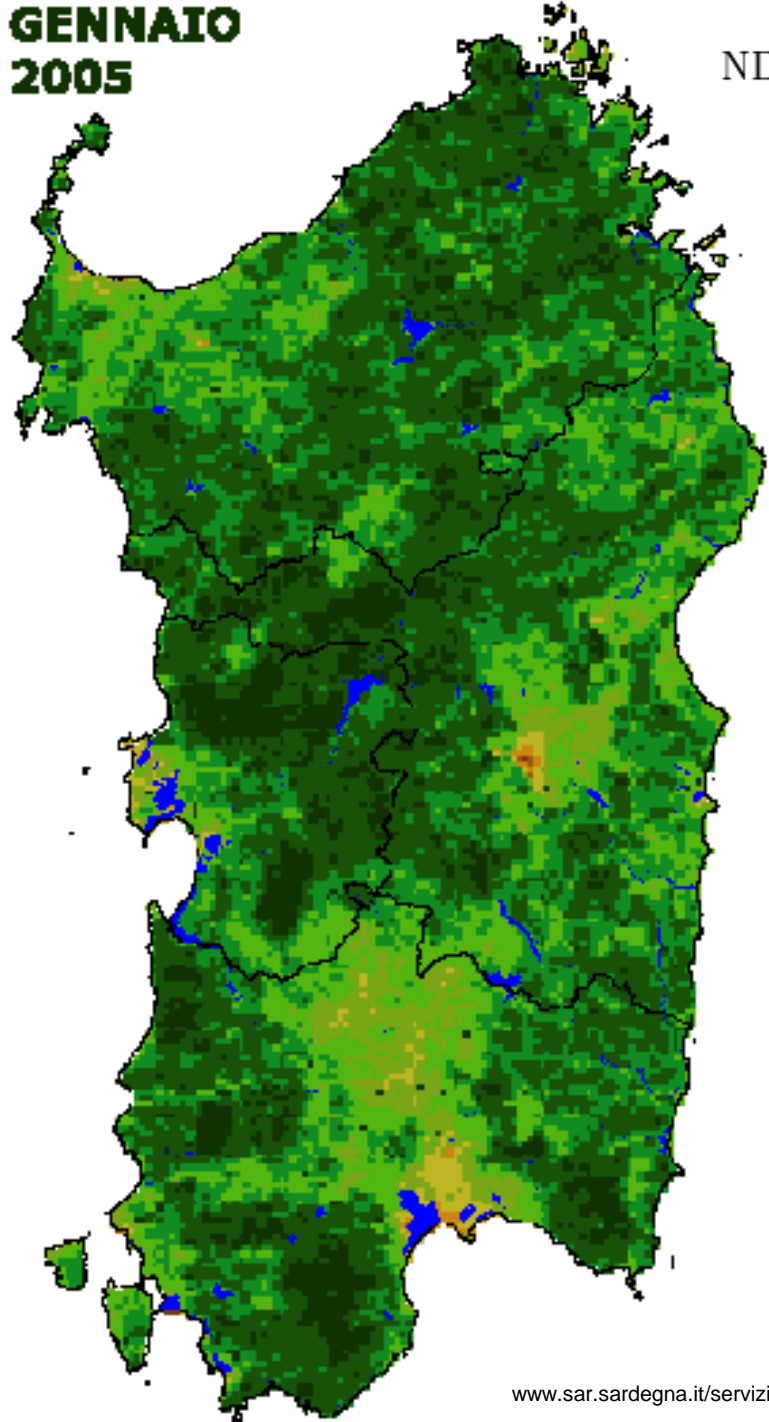
**Ministero
dell’Università
e della Ricerca**



Italiadomani
DIPARTIMENTO
DIPARTIMENTO

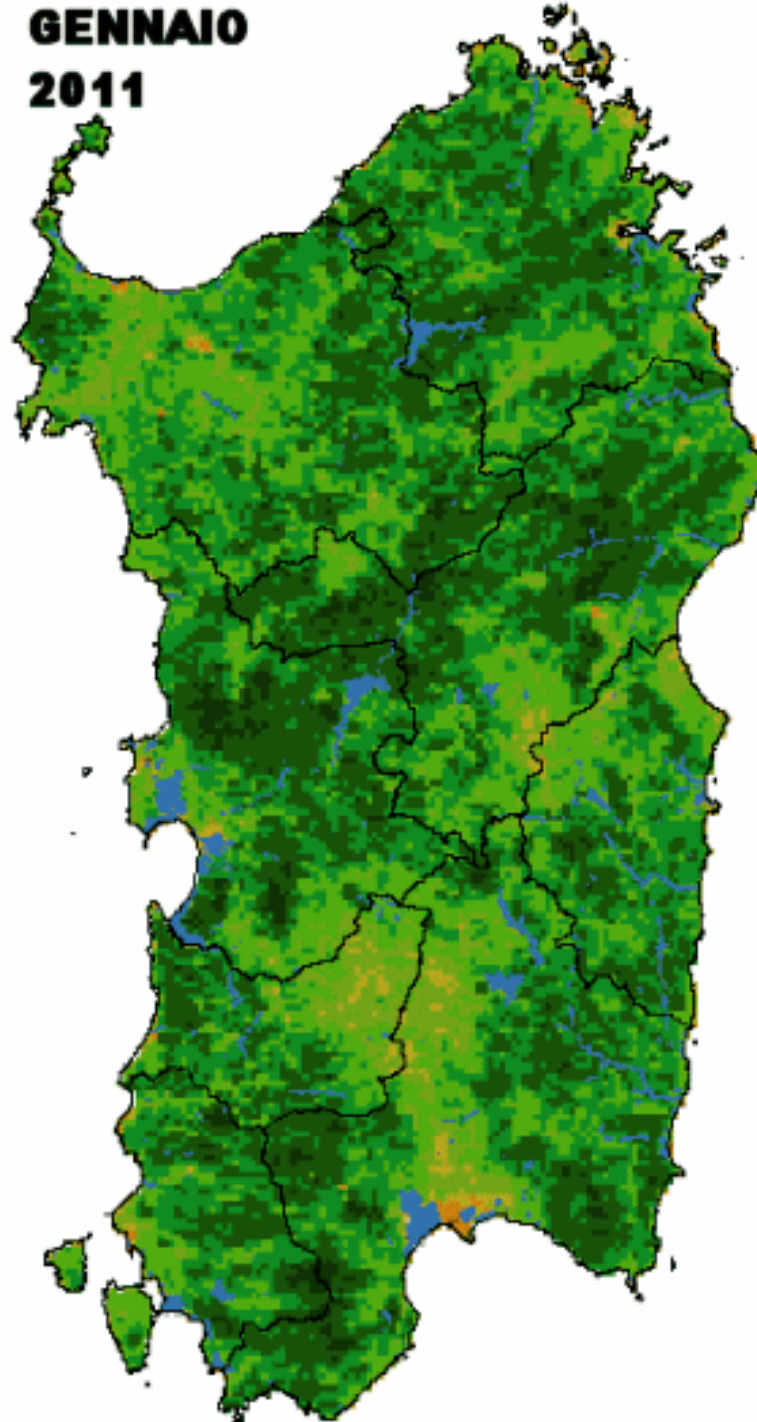


**GENNAIO
2005**

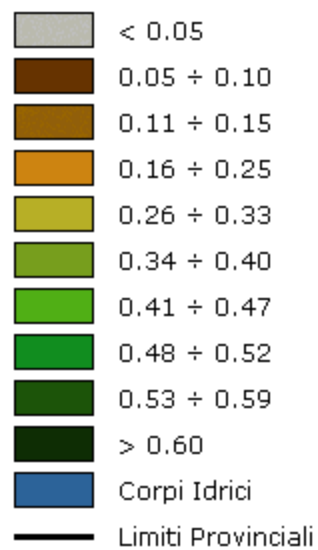


$$NDVI = \frac{(NIR - VIS)}{(NIR + VIS)}$$

**GENNAIO
2011**



Legenda



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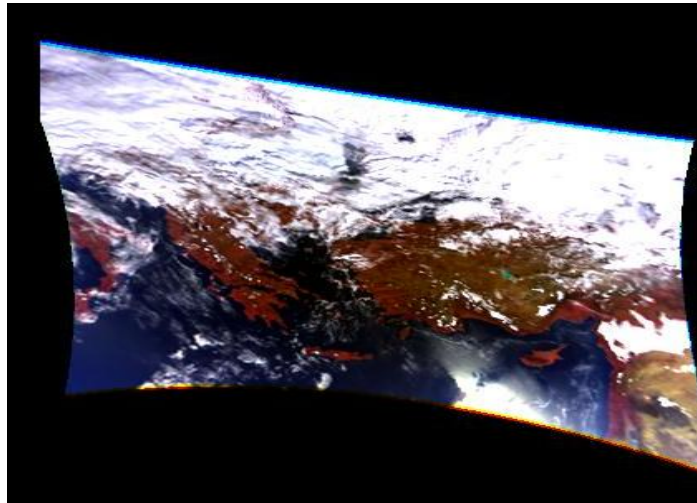
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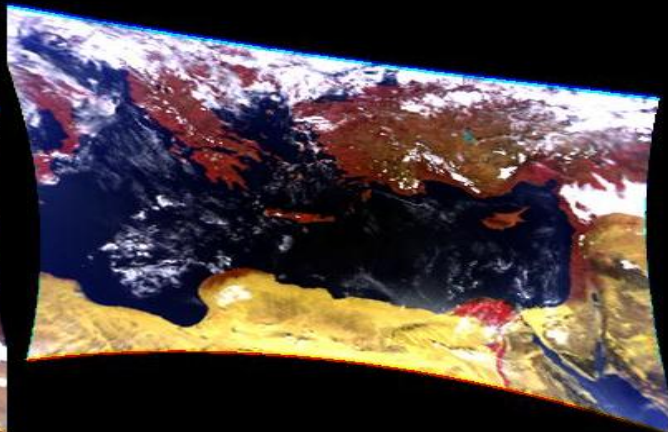
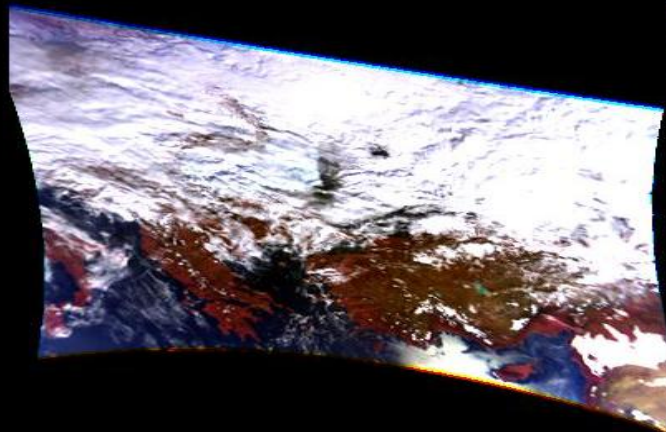
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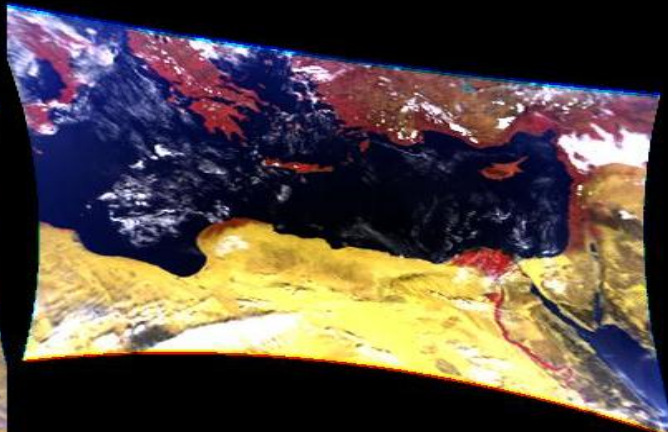
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White=Cloud

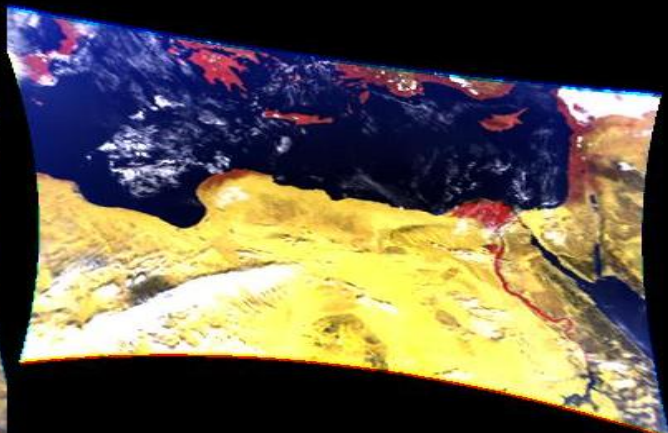




Cloud



Sunlint



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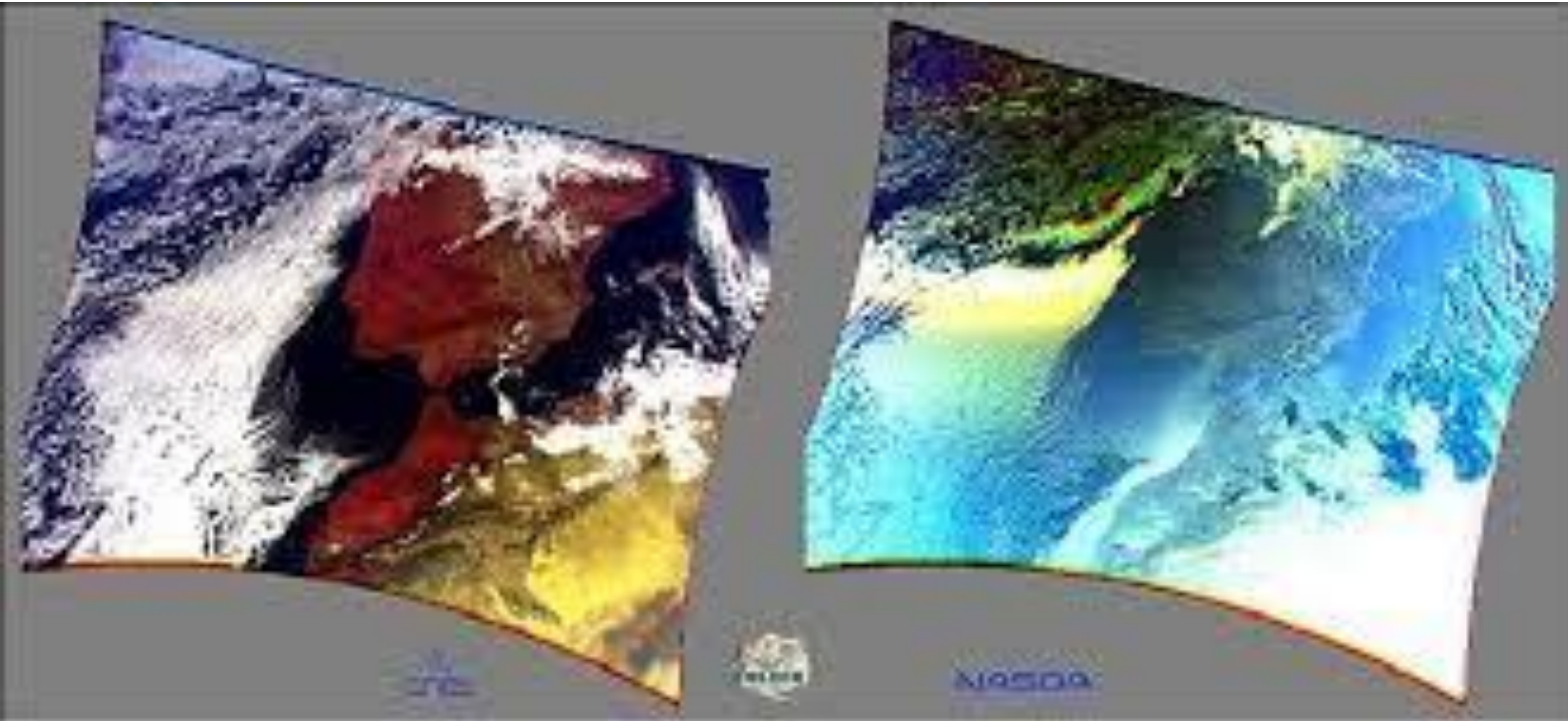
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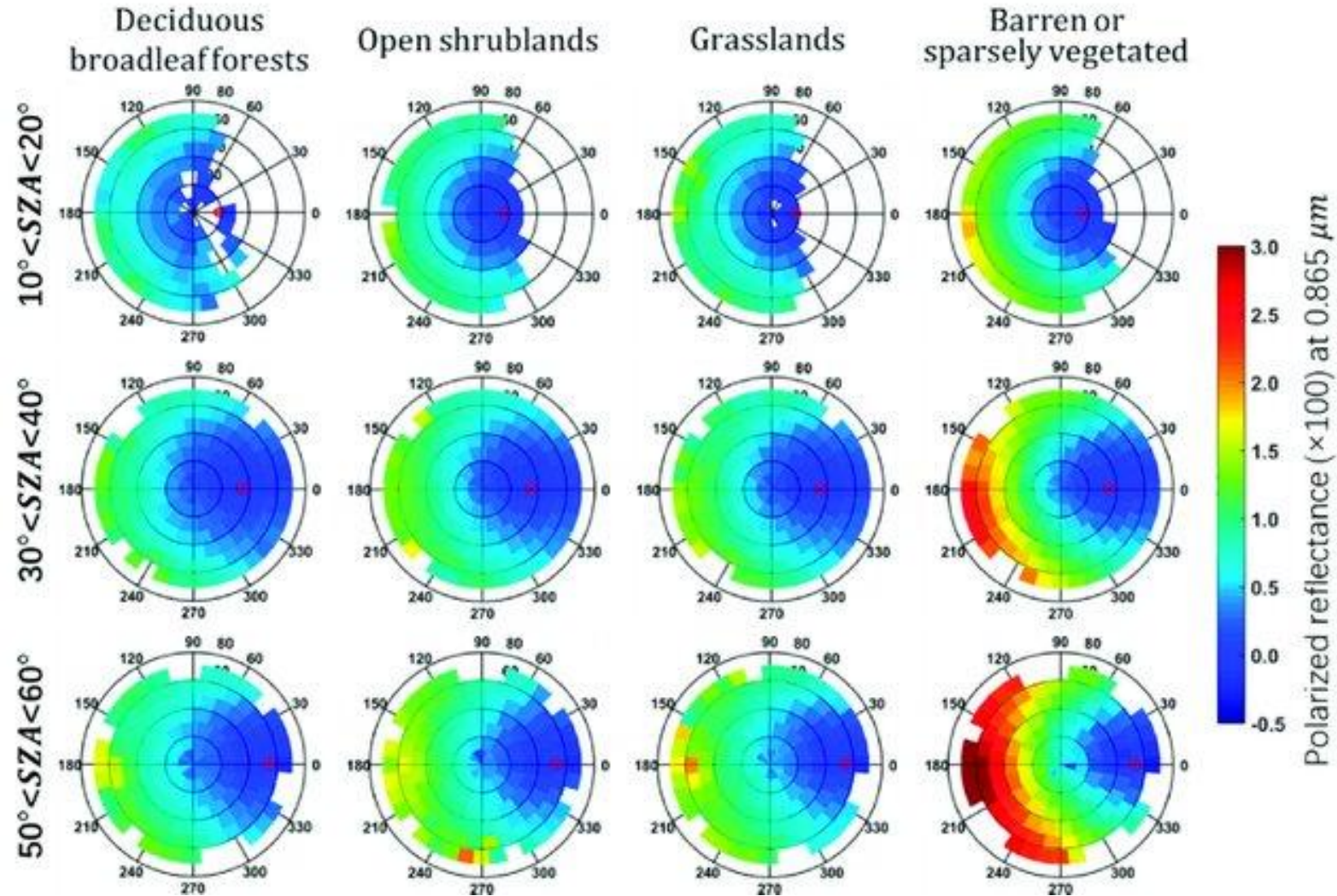
Example of POLDER RGB composites

Intensity

Polarization



Example of polarized reflectance directional distribution as a function of vegetation type



Data Inversion Techniques: Outline

🌐 Information 'Space' dimensions

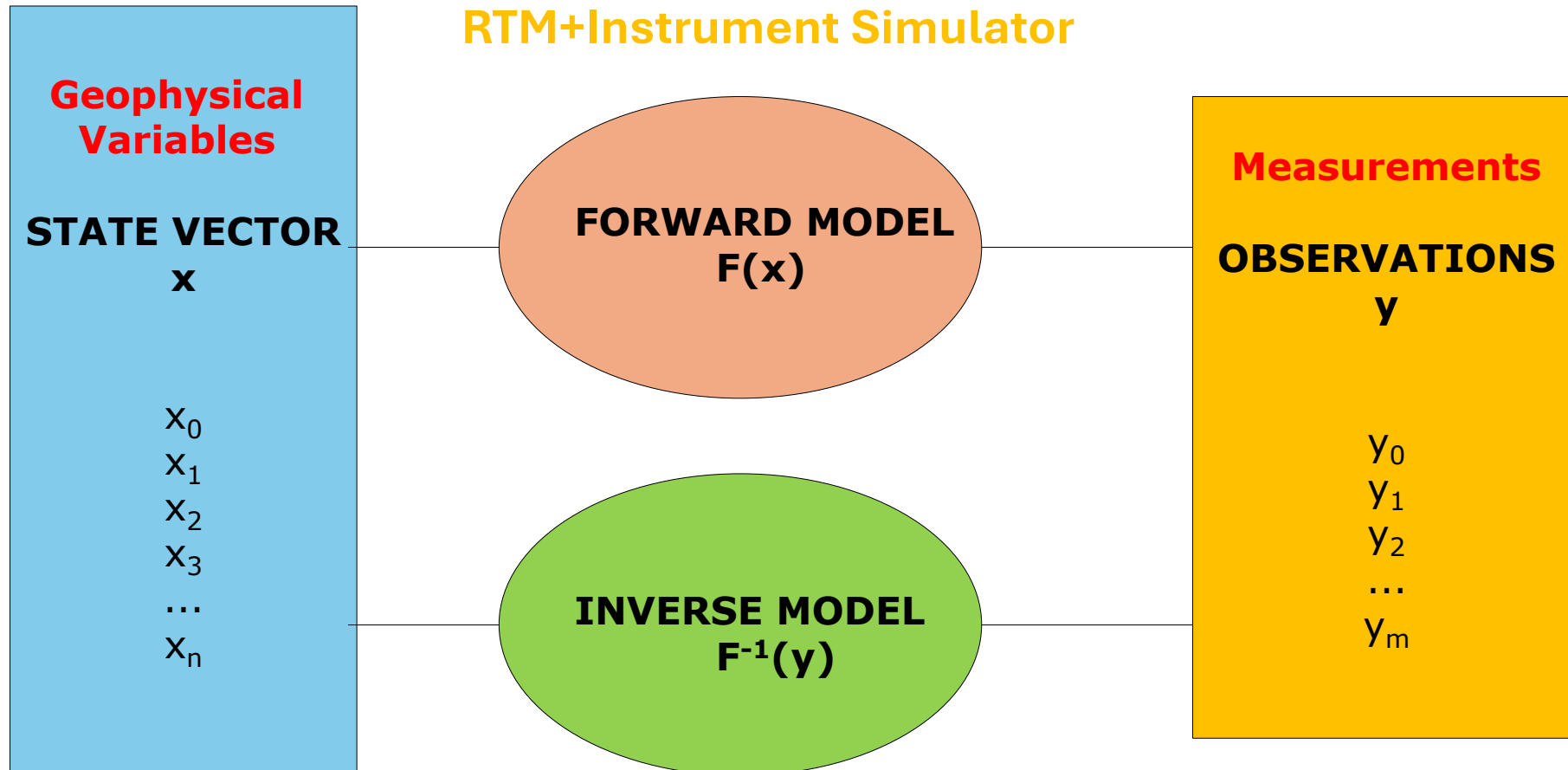
🌐 **Inversion Approaches**

Inversion Approaches: Ingredients

- 🌐 **X**: *State Vector*: vector, of dimension n , containing all relevant variables, among which is the searched solution, from which the measurements can depend.
- 🌐 **Y**: *Measurement Vector*: Vector of dimension m , containing all relevant measurements.
- 🌐 ϵ_Y : *Measurement Uncertainty Vector/array*: Vector/array of dimension $m/m \times m$, containing the uncertainties of all relevant measurements (when array includes also the covariance information).
- 🌐 **A**: *Ancillary data*: set of *known/assumed* informations relevant to solve the problem: geographical position, time, observation/illumination geometry, relevant geophysical variables from external data sources (e.g. climatology, forecast model).
- 🌐 **F**: Forward model. A model that given a geophysical scenario and acquisition conditions (geometry) gives the expected measurement vector: $F(X)=Y$

Forward problem

From geophysical variables to measurements



Inverse problem

From measurement to geophysical variables

A non trivial problem

- Forward model F very often not linear and not explicitly invertible: .
- Measurements do not depend **uniquely** from the variable of interest
- *Hill posed* problems: different number of observations (known) and unknown. **Possibility of different combination of solutions that give the same result in terms of measurements.**
- Uncertainties in the measurements and in the forward model

Inversion- Possible approaches

- **Analytical:** not always possible: signal equation(s) should be analytically inverted for all relevant variable. Generally, need assumptions/ancillary data to handle second order effects. *An estimation of the uncertainty may be obtained from uncertainty propagation laws.*
- **Variational** – does not requires assumptions on the set of possible solutions (may limit their range). Numerically inefficient. *Very likely the method gives an estimation of the uncertainty.*
- **Regressions (including Artificial Intelligence and similar):** Computationally efficient. Need a training set. Assumptions on the functional form/structure. May have difficulties in finding a solutions for cases not included in the training data set. *Uncertainty estimation may not be straightforward.*
- **Lookup Table (LUT)** – computational efficiency dependent from the amount of LUTs, Search the solution within a set of tabulated (LUT) solutions. Set of possible solutions defined (difficulties in extrapolating). *Difficult to estimate the uncertainty of the solution*
- **Index (e.g. NDVI, AI)** – Relatively simple because defined from the measurements, but it does not represent a geophysical variable. It needs, for some application, a relationship with geophysical variables.

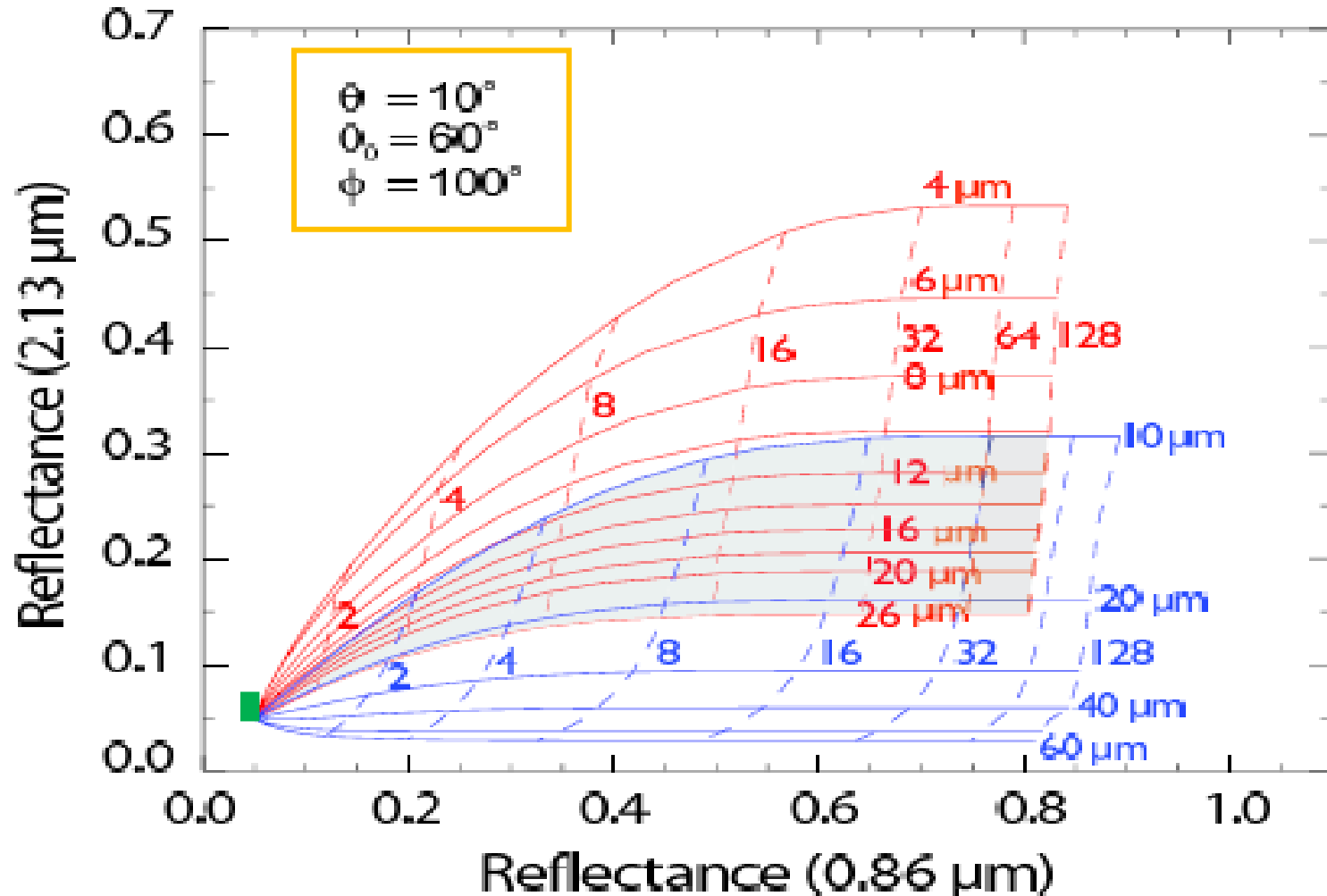
Example of LUT-based retrieval: cloud optical thickness & particle effective radius.

Inputs to select the LUT:

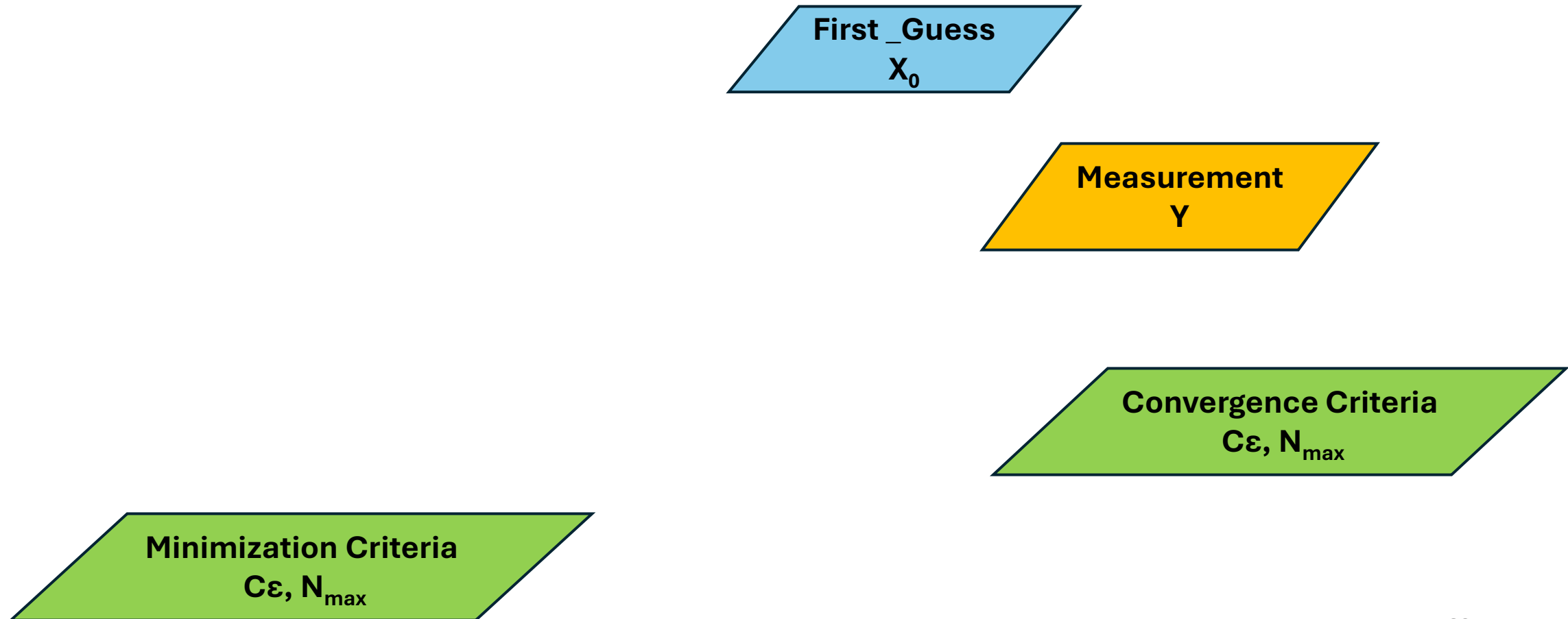
observation/illumination geometry, cloud top phase (liquid/ice), surface reflectance

Inputs to the LUT:

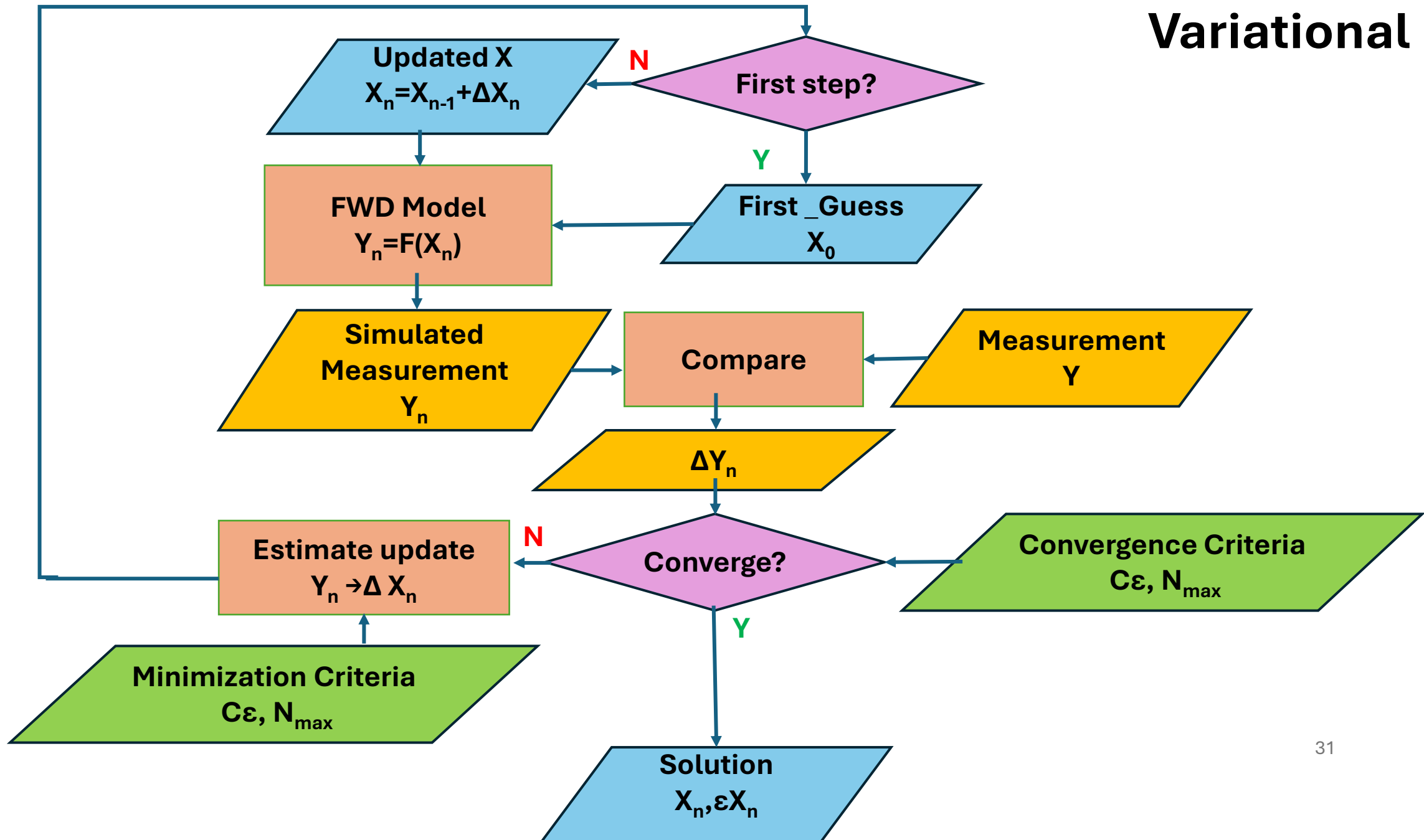
TOA reflectance @0.86 and @2.13



Variational

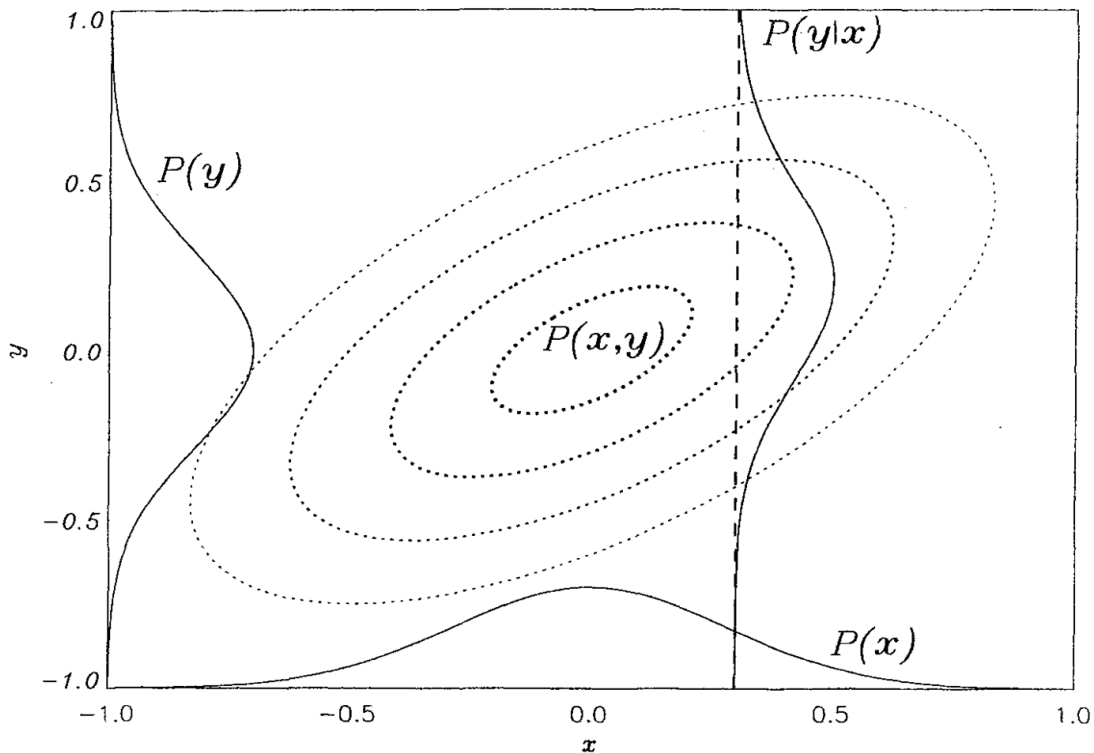


Variational



Bayesian Interpretation

The problem is set as maximizing the probability that, given a geophysical scenario characterized by a *state vector* \mathbf{x} (for ex: surface wind, temperature and water vapour amount profiles, cloud liquid water, sea surface temperature) the instrument measures the *measurement vector* \mathbf{y}



Joint Probability Conditional probability

$$P(x, y) = P(x | y)P(y)$$

$$P(y, x) = P(y | x)P(x)$$

$$P(x, y) = P(y, x)$$

Bayesian Interpretation

PDF of the measurement given a certain Status (x) includes the uncertainties due to the instrument precision

Informazione a priori sullo stato.
Probabilità che ci sia un dato stato indipendentemente dalla misura (es. climatologia)

BAYES THEOREM:

$$P(x | y) = \frac{P(y | x)P(x)}{P(y)}$$

Normalization factor
(a priori PDF of the measurement)
 $P(y)=1$

- Possiamo assumere una incertezza sperimentale con distribuzione gaussiana

$$P(y | x) = A \exp \left(- \frac{(y - F(x))^2}{2\sigma^2} \right)$$

- Per il caso multidimensionale abbiamo invece

$$P(y | x) = A \exp \left(- \frac{1}{2} (y - F(x))^T C_\epsilon^{-1} (y - F(x)) \right)$$

- Ragionevole assumere una distribuzione gaussiana anche per l'informazione a priori su x

$$P(x) = A \exp \left(- \frac{1}{2} (x - x_a)^T C_a^{-1} (x - x_a) \right)$$

- Otteniamo la seguente probabilità a posteriori

$$P(x | y) = A \exp\left(-\frac{1}{2}\left((y - F(x))^T C_\epsilon^{-1}(y - F(x)) + (x - x_a)^T C_a^{-1}(x - x_a)\right)\right)$$

- Noi vogliamo una soluzione, non l'intera PDF.
Dobbiamo scegliere cosa ottimizzare, es: il valore più probabile, minimizzare l'incertezza, massimizzare la risoluzione

- MAP (Maximum A posteriori Probability)

$$x = \mathit{max}P(x | y)$$

- Il massimo può essere calcolato in vari modi:
 - - **Analiticamente** (se possibile)
 - - **Ottimizzazione numerica** for ex.: Gauss-Newton, Gradient coniugato
 - - **Monte Carlo**

- Nel caso **non sia presente una conoscenza a priori** di x , la soluzione MAP è la stessa della ML (Maximum likelihood)

$$P(y | x) = A \exp\left(-\frac{1}{2}(y - F(x))^T C_\epsilon^{-1}(y - F(x))\right)$$

- LSM (Least square method)

Cost function $\Psi(x) = \frac{1}{2}(y - F(x))^T C_\epsilon^{-1}(y - F(x)) = \min$

$$\nabla \Psi(x) = \frac{\partial \Psi(x)}{\partial x_i} = 0$$

• Maximum likelihood and LSM

For a **linear*** system for which $\mathbf{f}(\mathbf{x})=\mathbf{K}\mathbf{x}$ we can write the Cost Function $\Psi(x)$ as:

$$\Psi(x) = \frac{1}{2} (y - Kx)^T C^{-1} (y - Kx)$$

Minimising $\Psi(x)$ gives:

$$\nabla\Psi(x) = K^T C^{-1} Kx - K^T C^{-1} y = 0$$

↓

$$K^T C^{-1} Kx = K^T C^{-1} y$$

Solving in x :

$$x = (K^T C^{-1} K)^{-1} K^T C^{-1} y$$

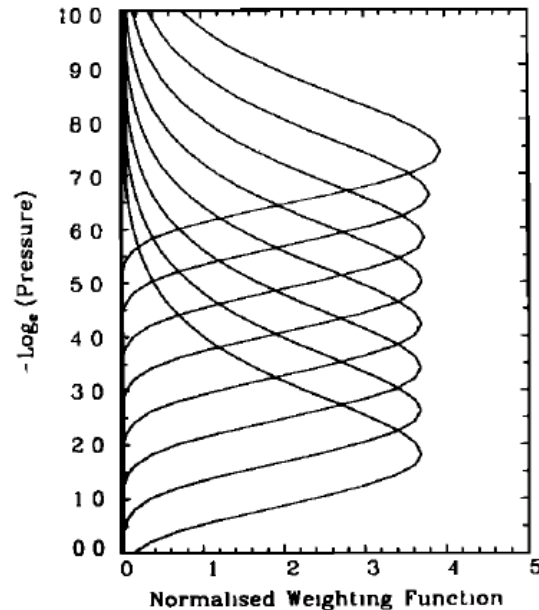
Valid if $\det(K^T C^{-1} K) \neq 0$ and uncertainties have a gaussian distribution

• * It is possible to linearize a non linear operator $f(x)$ in a given interval

Regularizzazione

- Nel caso di problemi mal posti non posso invertire K

$$y = Kx \quad \longrightarrow \quad \begin{pmatrix} y_1 \\ y_2 \\ \vdots \\ y_m \end{pmatrix} = \begin{pmatrix} k_{11} & k_{12} & \cdots & k_{1n} \\ k_{21} & k_{22} & \cdots & k_{2n} \\ k_{31} & k_{32} & \cdots & k_{3n} \\ \vdots & \vdots & \ddots & \vdots \\ k_{m1} & k_{m2} & \cdots & k_{mn} \end{pmatrix} \begin{pmatrix} x_1 \\ x_2 \\ x_3 \\ \vdots \\ x_n \end{pmatrix}$$



• A set of idealized weighting functions normalized to unit area.

$$I(\nu) = \int_0^{\infty} B(\nu, z) \frac{dT}{dz}(\nu, z) dz$$

$$I_i = I(\nu_i) = \int_0^{\infty} B(\bar{\nu}, z) K_i(z) dz \quad i = 1 \cdots M$$

$$B(\bar{\nu}, z) = \sum_{j=1}^M b_j W_j(z)$$

$$I_i = \sum_{j=1}^M b_j \int_0^{\infty} W_j(z) K_i(z) dz = \sum_1^M A_{i,j} b_j$$

Regolarizzazione

- Nel caso di problemi mal posti non posso invertire K

$$y = Kx \quad \longrightarrow \quad \begin{pmatrix} y_1 \\ y_2 \\ \vdots \\ y_m \end{pmatrix} = \begin{pmatrix} k_{11} & k_{12} & \cdots & k_{1n} \\ k_{21} & k_{22} & \cdots & k_{2n} \\ k_{31} & k_{32} & \cdots & k_{3n} \\ \vdots & \vdots & \ddots & \vdots \\ k_{m1} & k_{m2} & \cdots & k_{mn} \end{pmatrix} \begin{pmatrix} x_1 \\ x_2 \\ x_3 \\ \vdots \\ x_n \end{pmatrix}$$

$$p = \text{rank}(K)$$

$$m = \text{dim}(y)$$

$$n = \text{dim}(x)$$

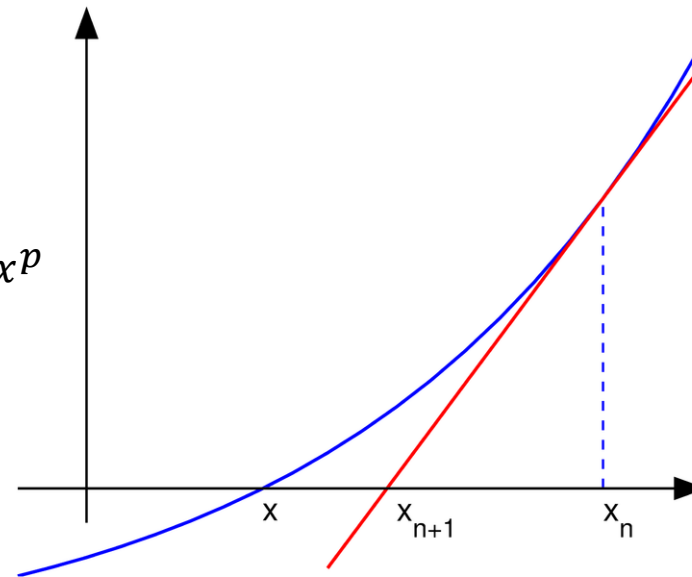
- **$p=m=n$** Well posed
- **$p < m=n$** Under and overconstrained
- **$p=m < n$** Underconstrained
- **$p < m < n$** Under and overconstrained
- **$p=n < m$** Overconstrained
- **$p < n < m$** Under and overconstrained

- In caso di funzioni non lineari possiamo risolvere l'equazione $\mathbf{y}=\mathbf{f}(\mathbf{x})$ attraverso iterazioni basate su approssimazioni lineari e partendo da un valore iniziale
- Metodo di Newton

$$x^{p+1} = x^p - \Delta x^p$$

$$K_p \Delta x^p \approx f(x^p) - y \quad K_p = \nabla f(x) |_{x^p}$$

$$x^{p+1} = x^p - K_p^{-1}(f(x^p) - y)$$



- Vogliamo risolvere
- Sostituiamo

$$\nabla\Psi(x) = 0$$

$$f(x^p) = \nabla\Psi(x^p) \quad y = 0$$

e

- $x^{p+1} = x^p - \Delta x^p$

Normal system

$$A_p \Delta x^p \approx \nabla\Psi(x^p)$$

$$A_p = \nabla(\nabla\Psi(x)) |_{x^p} \approx K^T C^{-1} K$$

$$x^{p+1} = x^p - (K_p^T C^{-1} K_p)^{-1} \nabla\Psi(x^p) =$$

$$= x^p - (K_p^T C^{-1} K_p)^{-1} K^T C^{-1} (f(x^p) - y)$$

Numerical method

Numerical inversion method: Newton-Gauss iteration algorithm (Rodgers, 2000)

$$x^{(n+1)} = x^n - (S_a^{-1} + J_n^T S_\epsilon^{-1} J_n)^{-1} (J_n^T S_\epsilon^{-1} (F(x^n) - y) + S_a^{-1} (x_n - x_a))$$

x_a First guess vector

y Measurement vector

$F(x_n)$ Radiances vector at TOA generated by x_n

J Jacobi matrix, J^T Transposed Jacobi matrix

S_ϵ Instrumental error Covariance matrix

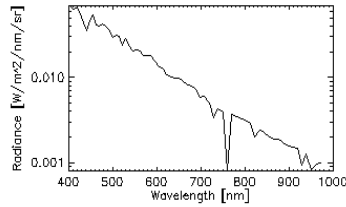
S_a First guess Covariance matrix

Example of numerical experiments

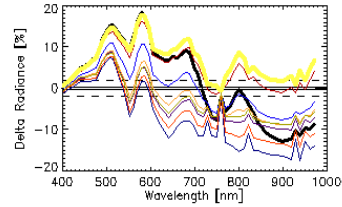
Spectra of % difference first guess, final solution

Simulated Spectrum
measurement vector

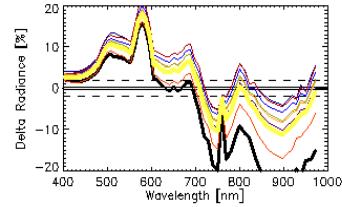
Exp.1



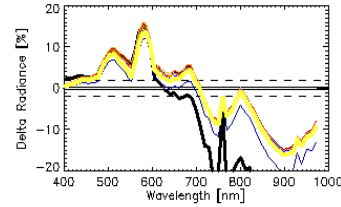
First guess 1



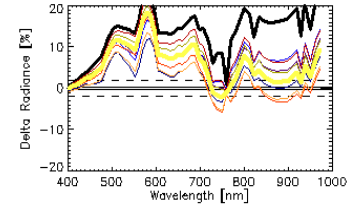
2



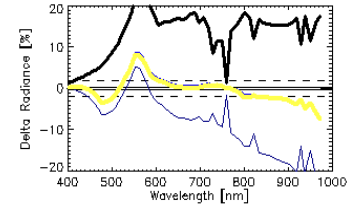
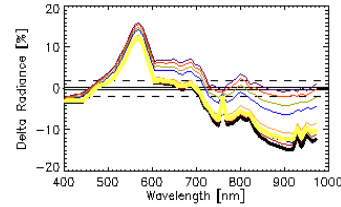
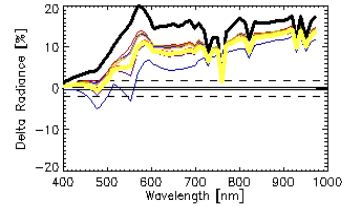
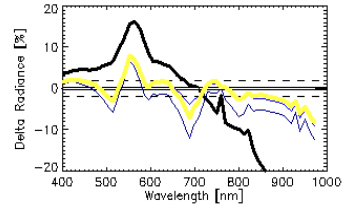
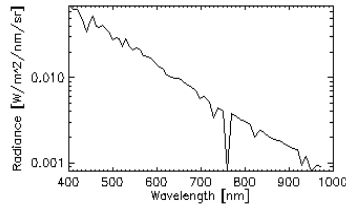
3



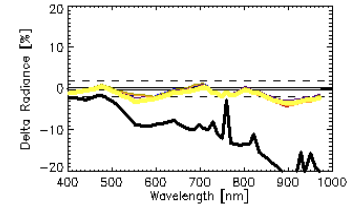
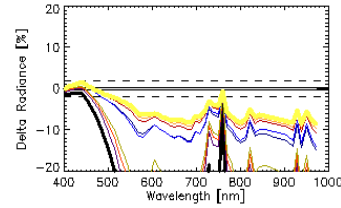
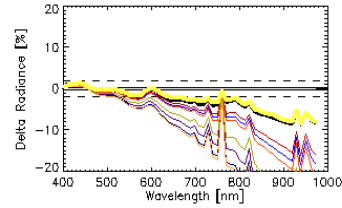
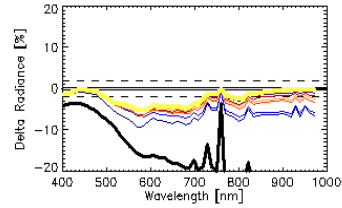
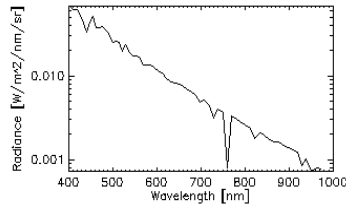
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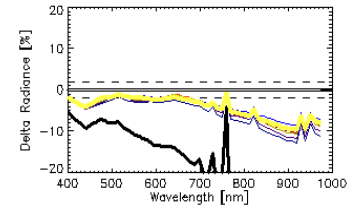
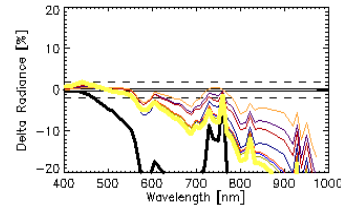
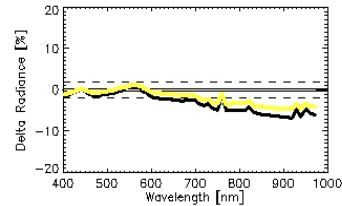
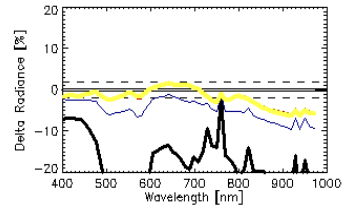
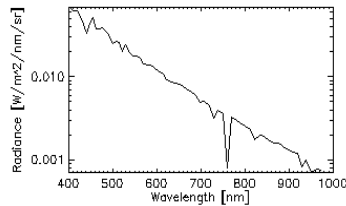
Exp.2



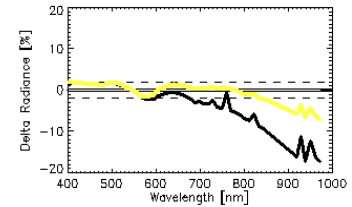
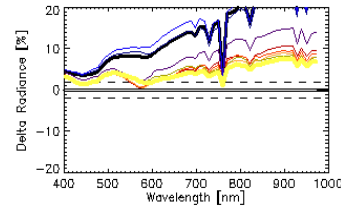
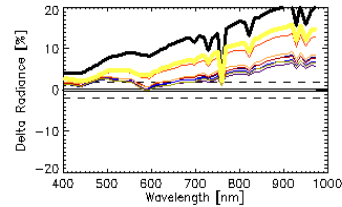
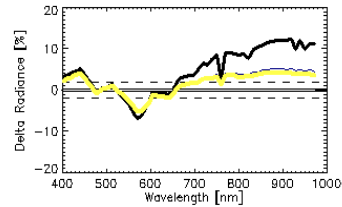
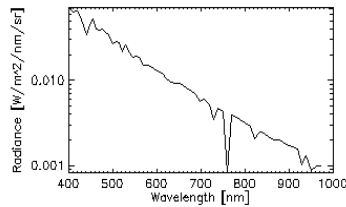
Exp.3



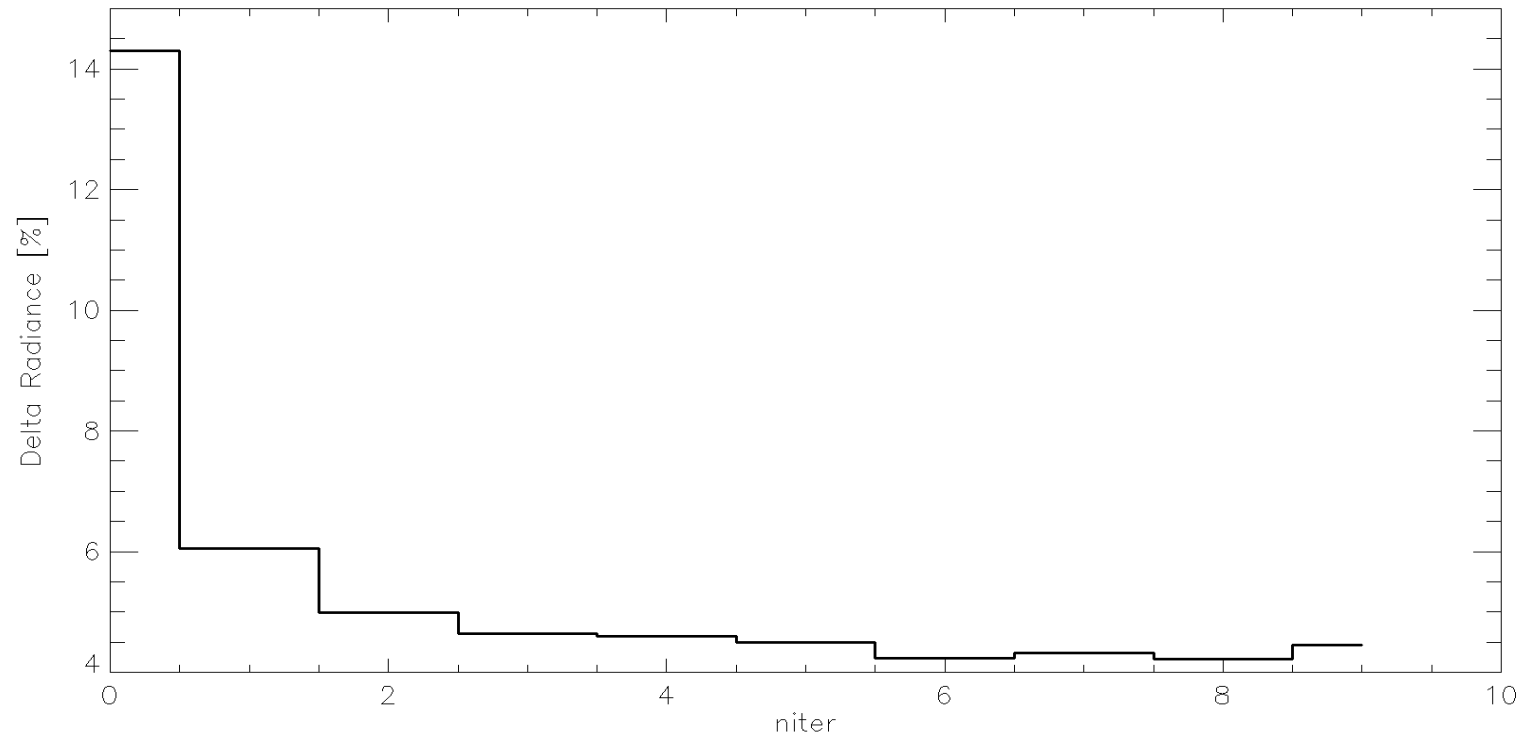
Exp.4



Exp.5

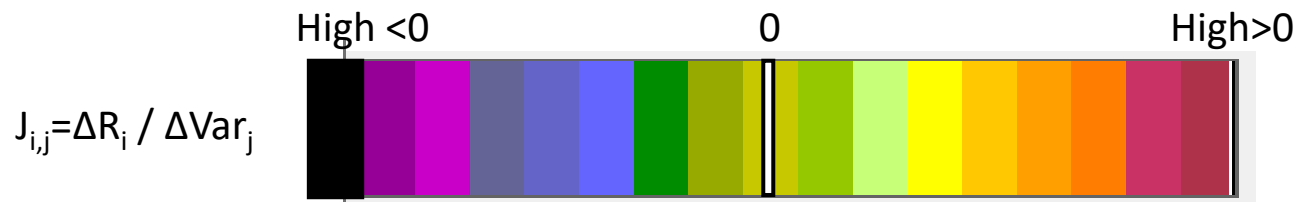


Average Δ Radiance [%] as a function of the iteration step



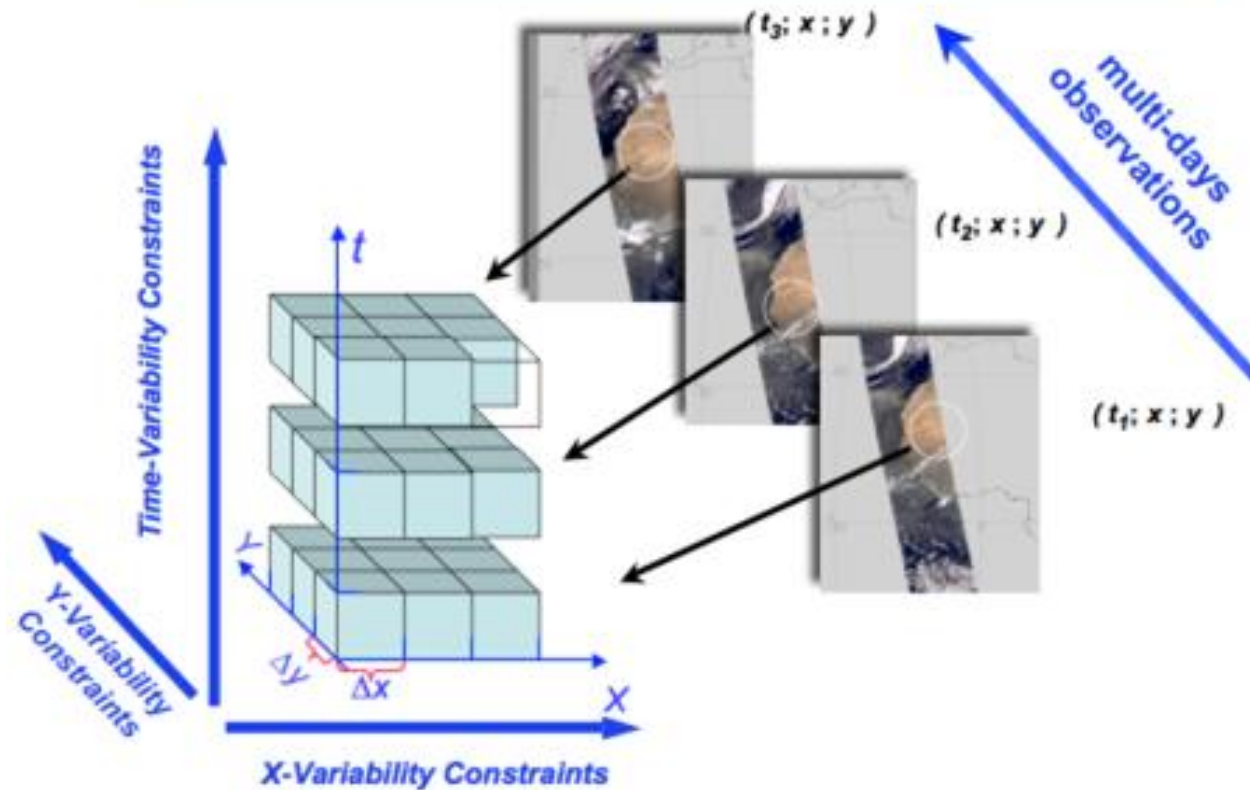
Example of Jacobian (average of 40 Jacobians)

18

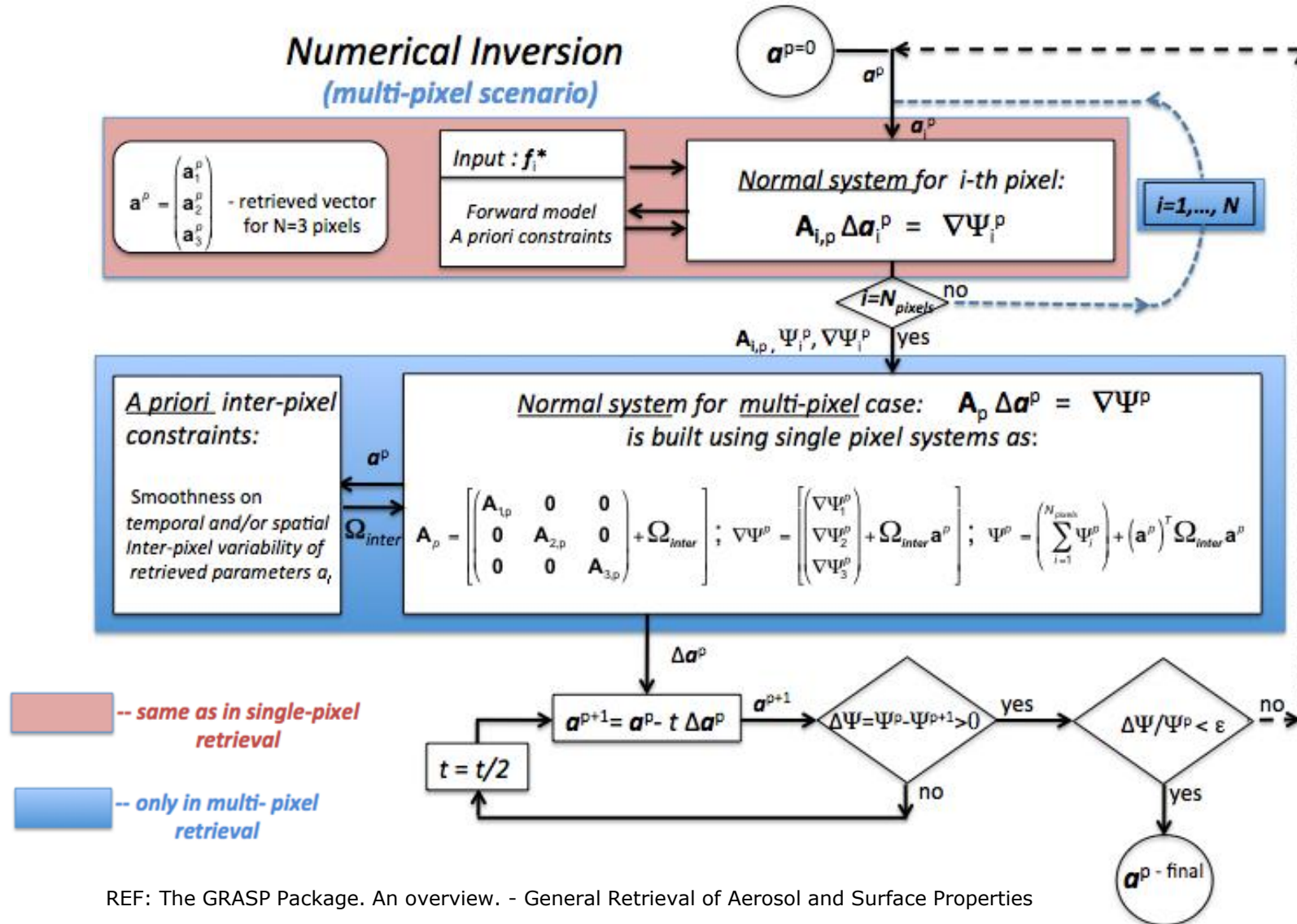


GRASP multi-pixel

The concept of multi-pixel retrieval



GRASP multi-pixel



Bibliography

- Inverse Methods for Atmospheric Sounding: Theory and Practice - Rodgers - 2000
- Inverse Problem Theory and Methods for Model Parameter Estimation - Tarantola - 2005
- Optimization of Numerical Inversion in Photopolarimetric Remote Sensing - Dubovik – 2005
- <https://www.grasp-open.com/products/aod-inversion/>
- <https://agupubs.onlinelibrary.wiley.com/doi/10.1029/RG014i004p00609>
- <https://www.ecmwf.int/en/elibrary/16942-inversion-methods-satellite-sounding-data>



THANKS!

IR0000032 – ITINERIS, Italian Integrated Environmental Research Infrastructures System
(D.D. n. 130/2022 - CUP B53C22002150006) Funded by EU - Next Generation EU PNRR-
Mission 4 "Education and Research" - Component 2: "From research to business" - Investment
3.1: "Fund for the realisation of an integrated system of research and innovation infrastructures"

