



Dark measurement test

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IR0000032 – ITINERIS, Italian Integrated Environmental Research Infrastructures System
(D.D. n. 130/2022 - CUP B53C22002150006) Funded by EU - Next Generation EU PNRR-
Mission 4 “Education and Research” - Component 2: “From research to business” - Investment
3.1: “Fund for the realisation of an integrated system of research and innovation infrastructures”



What is the Dark measurement test?

The **DARK measurement** is a quality assurance test used to **characterize the electronic response** of the lidar detection system **in the absence of any optical signal**

What is the Dark measurement test?

It is designed to highlight:

- **Electronic baseline and noise,**
- **Laser Synchronous* Spurious counts** in photon-counting channels,
- **Laser Synchronous* Systematic Distortions (LSSD)** in the analog signal:

** that appear in sync with the laser firing, even though no photons are detected.*

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These distortions are typically caused by:

- **Las** • **Crosstalk** between **laser triggering** and **acquisition electronics,**
- ana • **Inadequate grounding or shielding,**
- * *that* • **Induced signals** from high-voltage switching or pulse driver circuits.

Why is it important to perform the Dark test?

The PI needs to be able to **identify and subtract*** Laser-Synchronous Systematic Distortions (**LSSD**)

- subtraction is only possible if the distortion pattern is stable and repeatable.

The DARK measurement test is a **mandatory part of lidar QA protocols** (e.g. ACTRIS/EARLINET) because it provides critical **insights into the integrity and stability of the electronic detection chain**. Even though no optical signal is recorded during the test, the information it reveals is essential for ensuring **accurate, traceable, and artifact-free data**.

Freudenthaler, V., 2018

Why is it mandatory to submit the Dark test?

- for both analog and photon counting

1) Detection of Laser-Synchronous Systematic Distortions (LSSD)

These distortions **can mimic real atmospheric signals**, especially at short ranges.

2) Evaluation of Photon-Counting Channel Stability

Detects spurious artificial counts including:

- Dark counts,
- Afterpulses,
- Laser Synchronous Spurious Counts (LSSC)

Ensures that the counting electronics are not falsely reporting signals.

Freudenthaler, V., 2018

Why is it mandatory to submit the Dark test?

- for both analog and photon counting

3) Analog Channel Baseline Verification

Identifies offsets, electronic noise, and LSSD in analog signals.

4) Identifies the **Pre-trigger Region** (or far range region)

5) Instrument **Diagnostics** and Maintenance

Helps detect:

- Failing or drifting electronics,
- Improper shielding or grounding,
- Temperature-induced instabilities

Freudenthaler, V., 2018

How to perform the Dark test?

- including extended dark

based on CARS QA procedures

How to perform the Dark test?

- including extended dark

A. Ensure the system is fully powered and configured as for a normal measurement:

Same high voltage (HV) settings for PMTs/APDs,
Same acquisition settings, binning, and sampling rate.

(Let the system thermally stabilize before the test !)

B. Suppress Optical Signal

Close the telescope input to avoid external light (No photons should reach the detectors).

C. Start Acquisition

based on CARS QA procedures

How to perform the Dark test?

- including extended dark

Normal dark (an)

1. Let the laser and acquisition electronics reach **thermal stability**

- **20-30 minutes**,

2. **Collect dark signal**

- **15-30 minutes**,
- high temporal resolution: **10 seconds/profile**.

- Perform **before every measurement set** (an),
- For continuous: When switching between night-time and day-time (an).

Extended Dark (an + pc)

1. Let the laser and acquisition electronics reach **thermal stability**

- **20-30 minutes**,

2. **Collect dark signal**

- **< 30 minutes**,
- high temporal resolution: **10 seconds/profile**.

- **Submit only once and after each upgrade.**

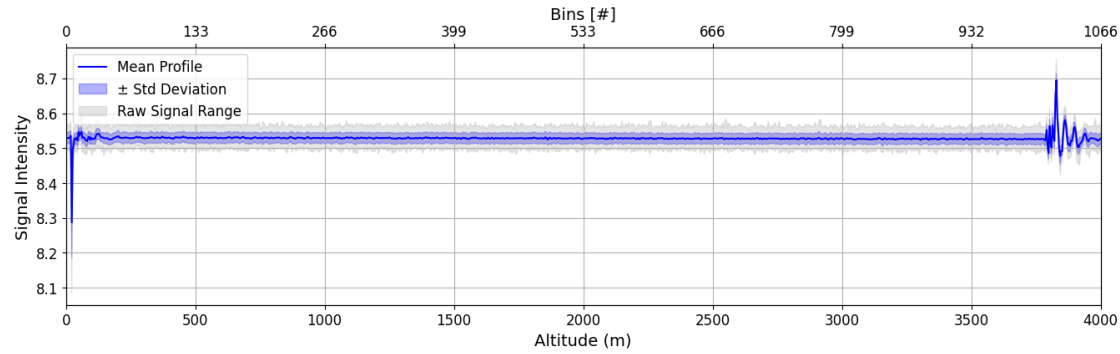
based on CARS QA procedures

Example Data

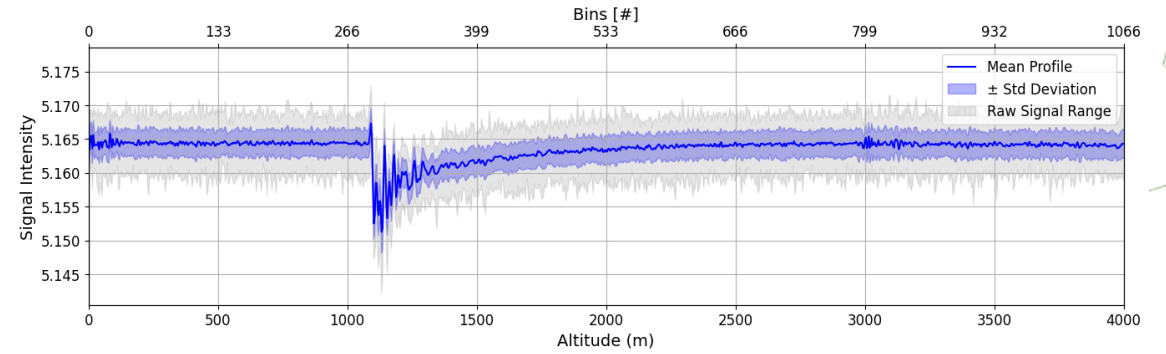
Assessment of pre-trigger region

analog

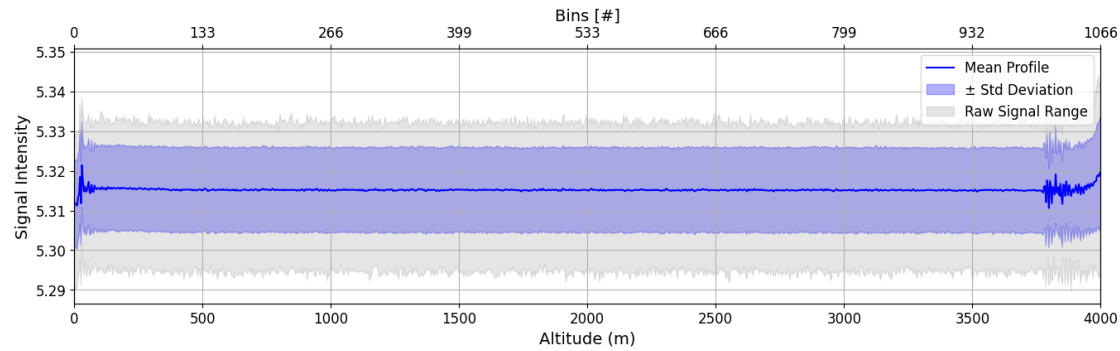
Dark Signal Profile - 1064.43 nm xtax
clj, CLOP, 20250221



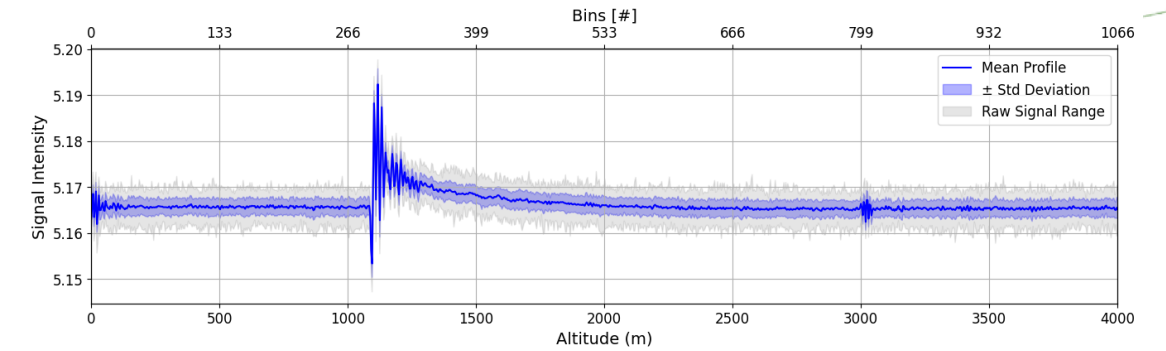
Dark Signal Profile - 532.2 nm fpar
pot, Fixed_Lidar_Raymetrics_L1, 20230928



Dark Signal Profile - 354.78 nm xpar
clj, CLOP, 20250221



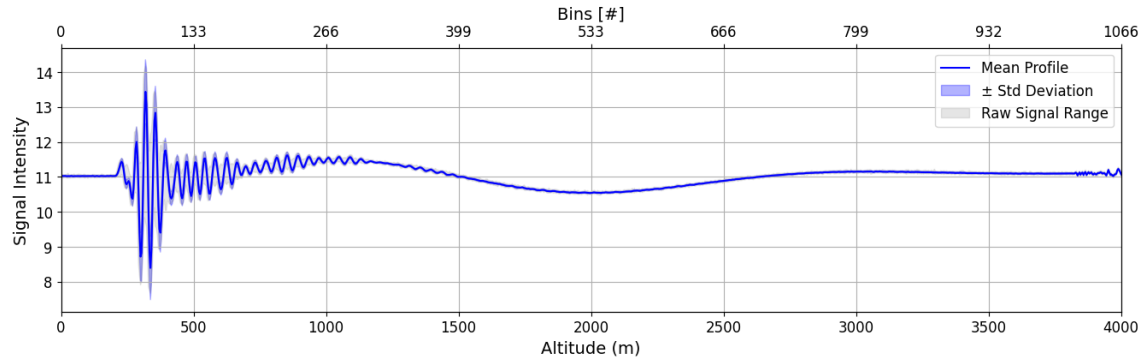
Dark Signal Profile - 532.2 nm fcat
pot, Fixed_Lidar_Raymetrics_L1, 20230928



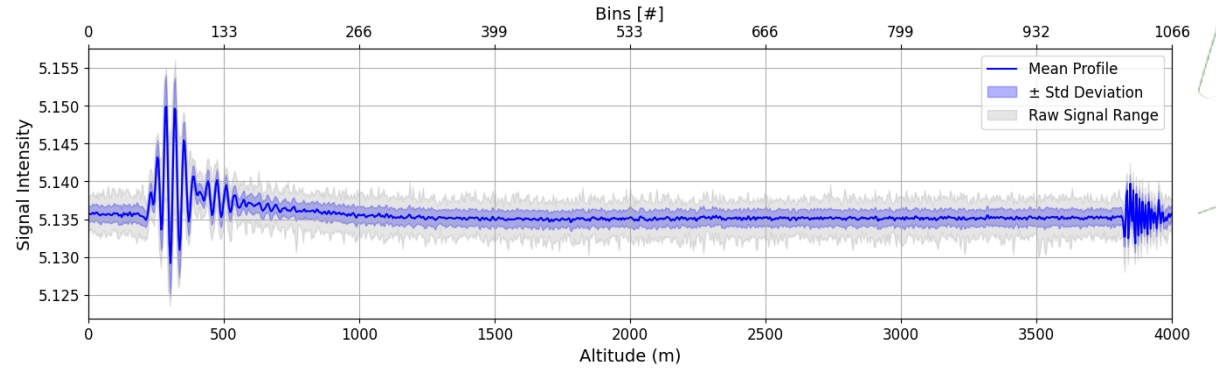
Assessment of pre-trigger region

analog

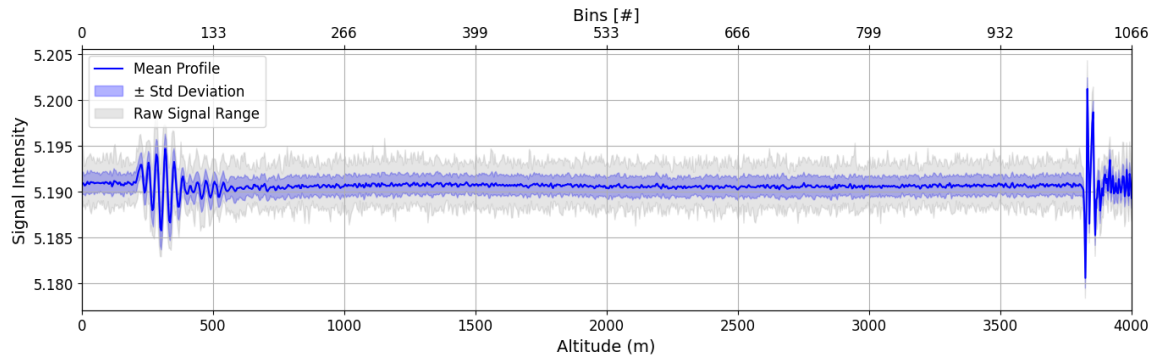
Dark Signal Profile - 1064.32 nm ftax
ino, A-lidar, 20240520



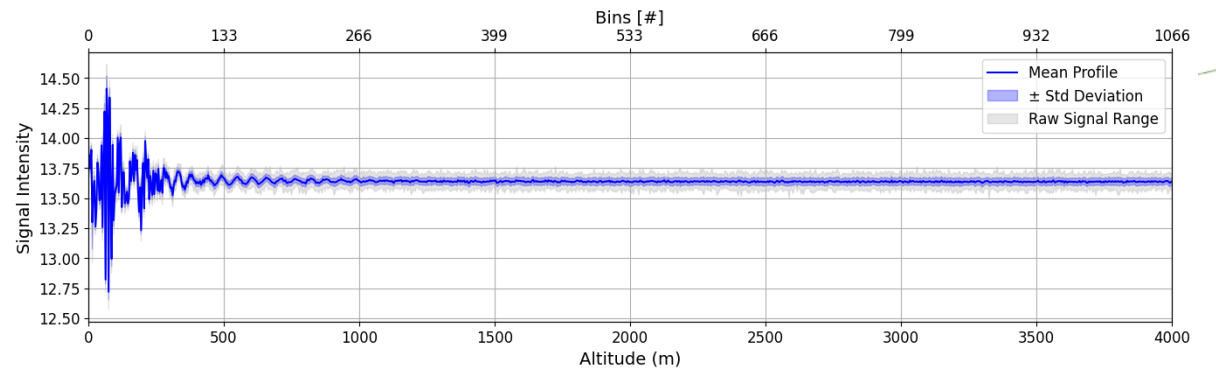
Dark Signal Profile - 354.0 nm fral
ino, A-lidar, 20240520



Dark Signal Profile - 354.9 nm ftax
ino, A-lidar, 20240520



Dark Signal Profile - 1067.43 nm xtax
rzc, ESA_Mobile_RAman_Lidar, 20240625

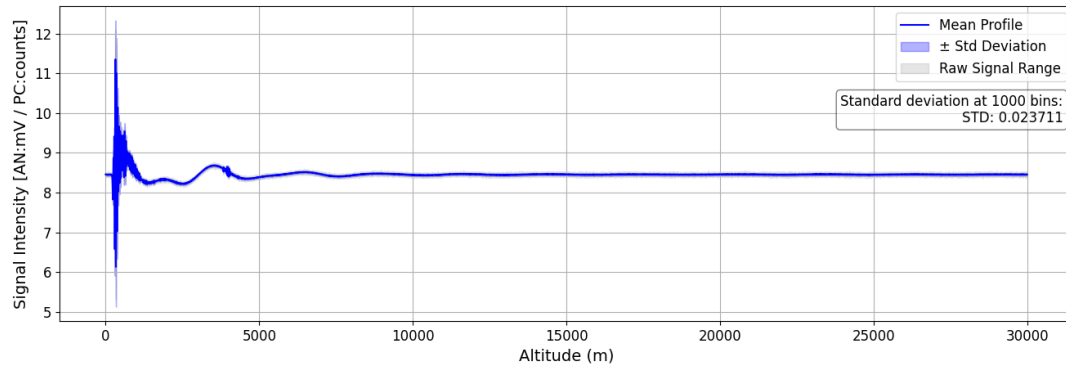


Assessment of dark signal

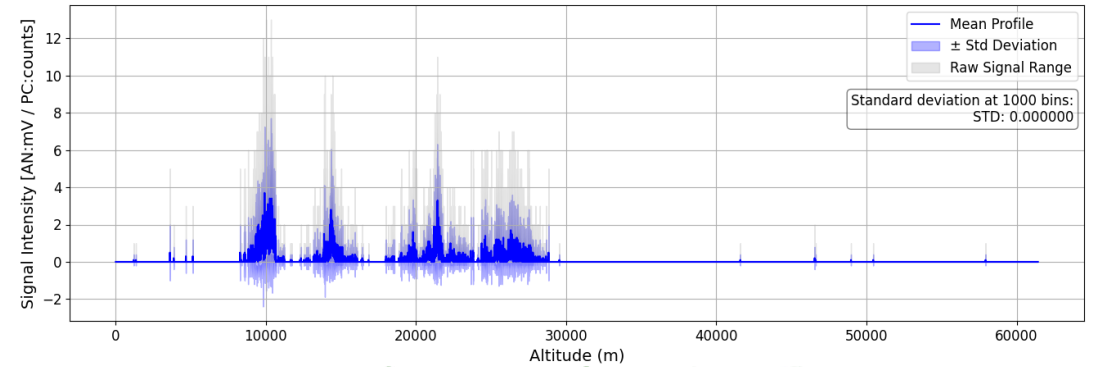
analog

photon-counting

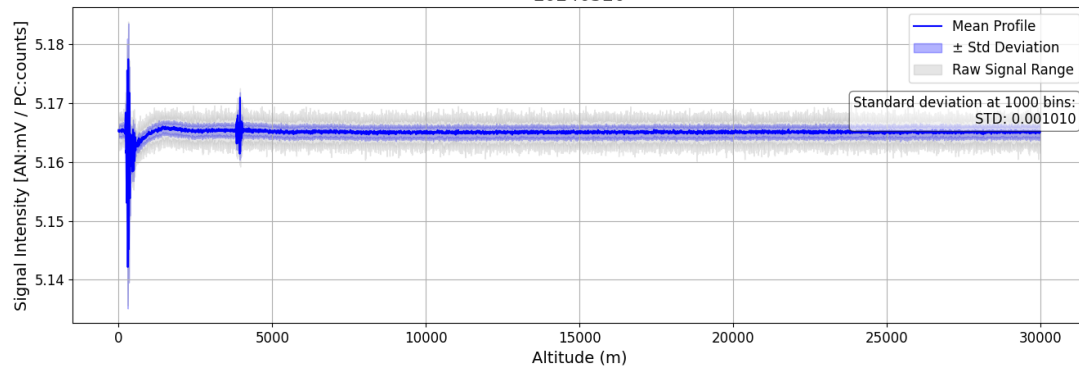
Dark Signal Profile - 1057.4 fral, ino, A-lidar
20240520



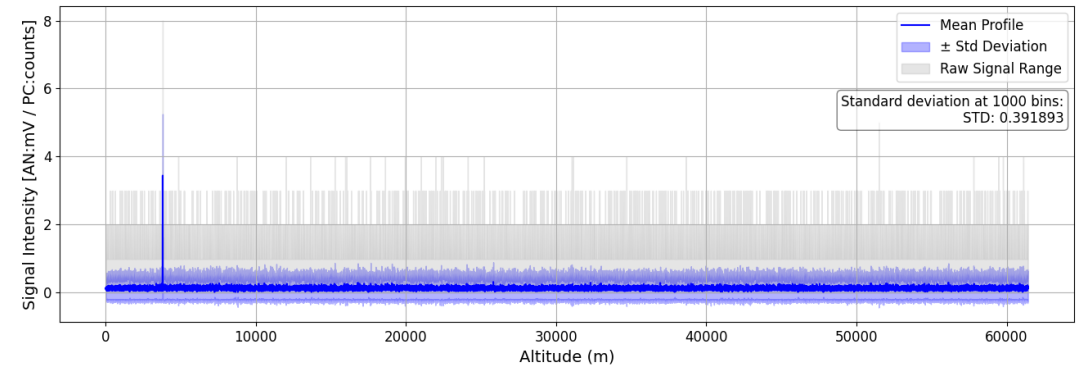
Dark Signal Profile - 532.26 xprr, rzc, ESA_MOBILE_Raman_Lidar
20240825



Dark Signal Profile - 532.2 ftax, ino, A-lidar
20240520



Dark Signal Profile - 532.26 xprr, clj, CLOP
20250225



General Recommendations

- Near-range interspersions can change, so it's necessary to **regularly test the temporal stability** of the dark signal for each channel before applying corrections.
- If the dark measurement time is **much shorter** than the atmospheric measurement (e.g., 10 min vs. 1:30 h), subtracting the dark measurement may **increase noise**, especially in the far range. **Smooth the Dark.... but**
- **Do not smooth** the dark signal **in the near range**; start smoothing only where signal noise increases.



THANKS!

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